

Exploring the Potential of Short-Baseline Physics at Fermilab



Pedro S. Pasquini
O. G. Miranda, M. Tórtola and J. W. F. Valle



06/19/2018

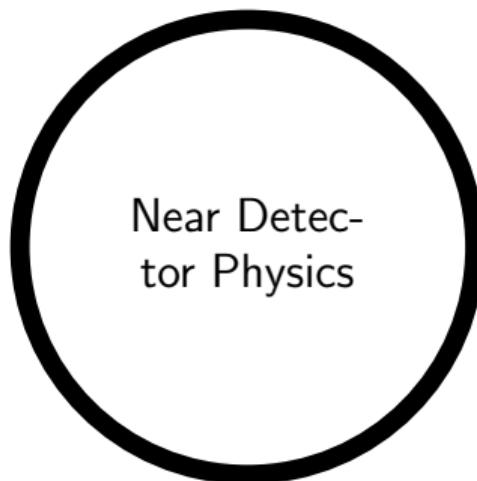
New Perspectives

Phys.Rev. D97 (2018) no.9, 095026

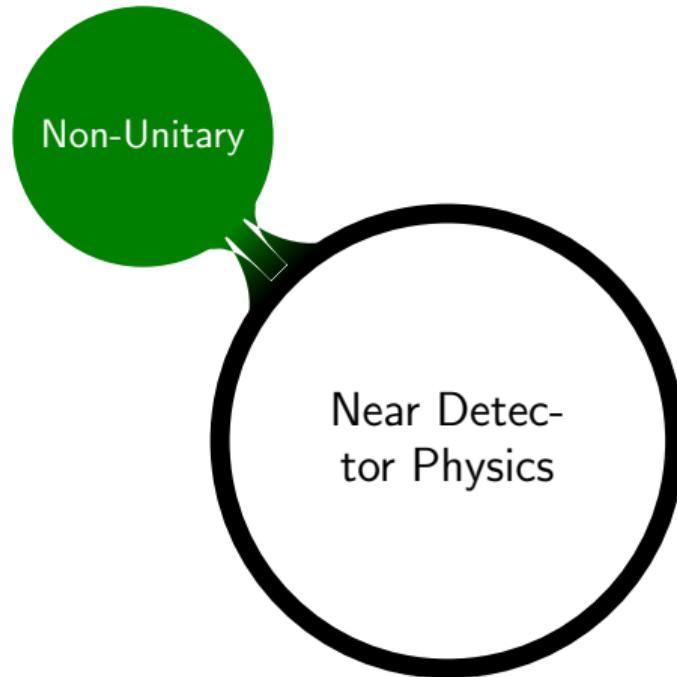
Arxiv: hep-pheno/1802.02133



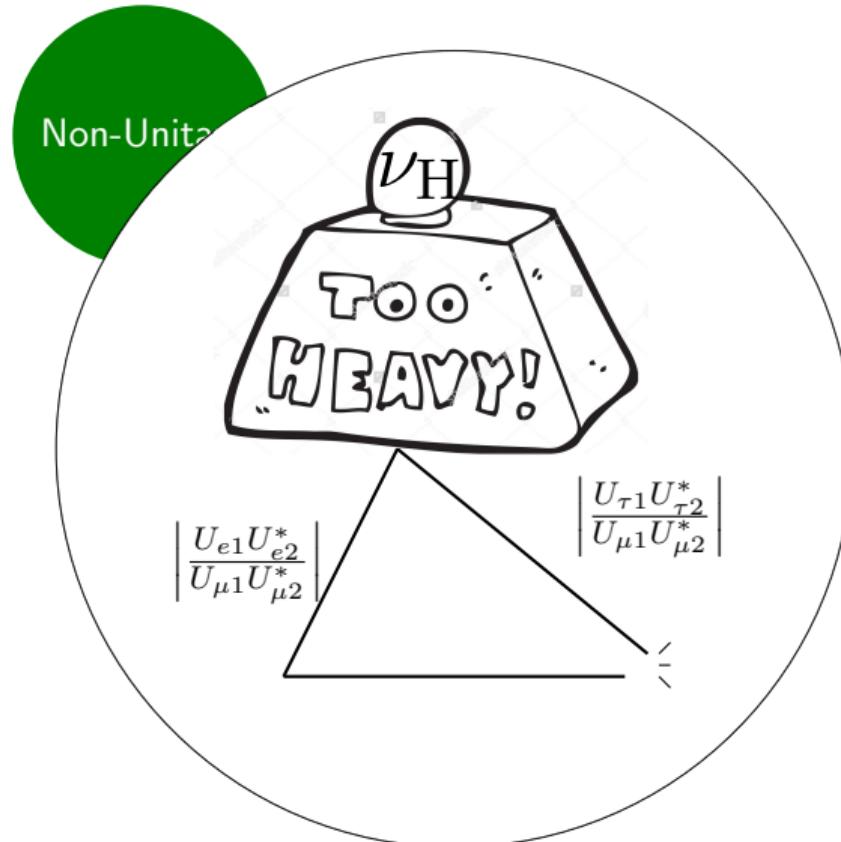
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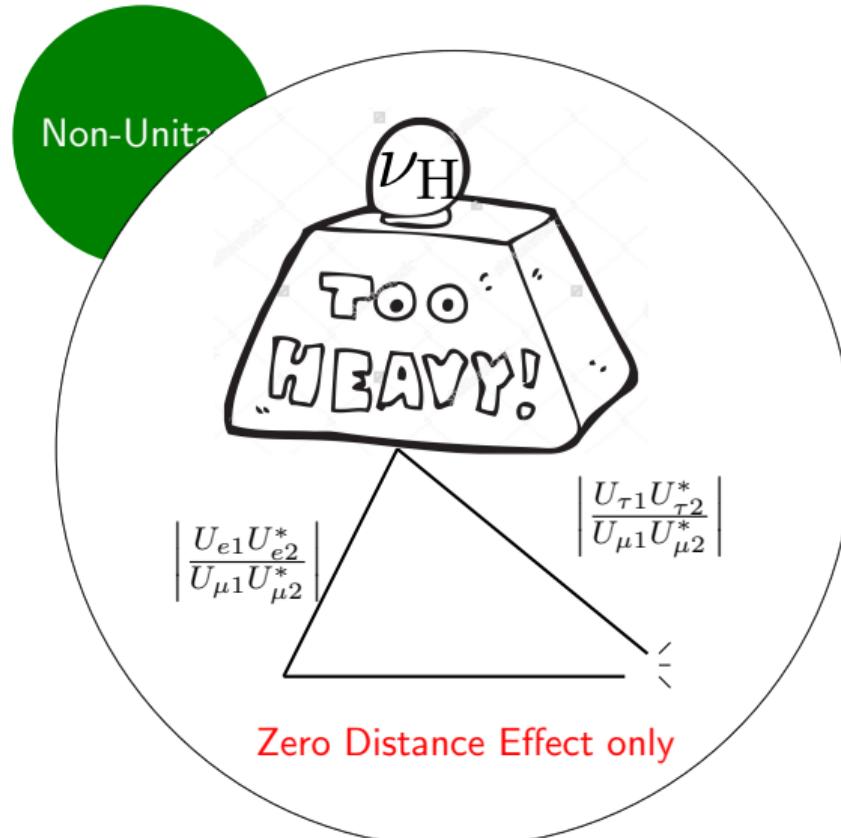
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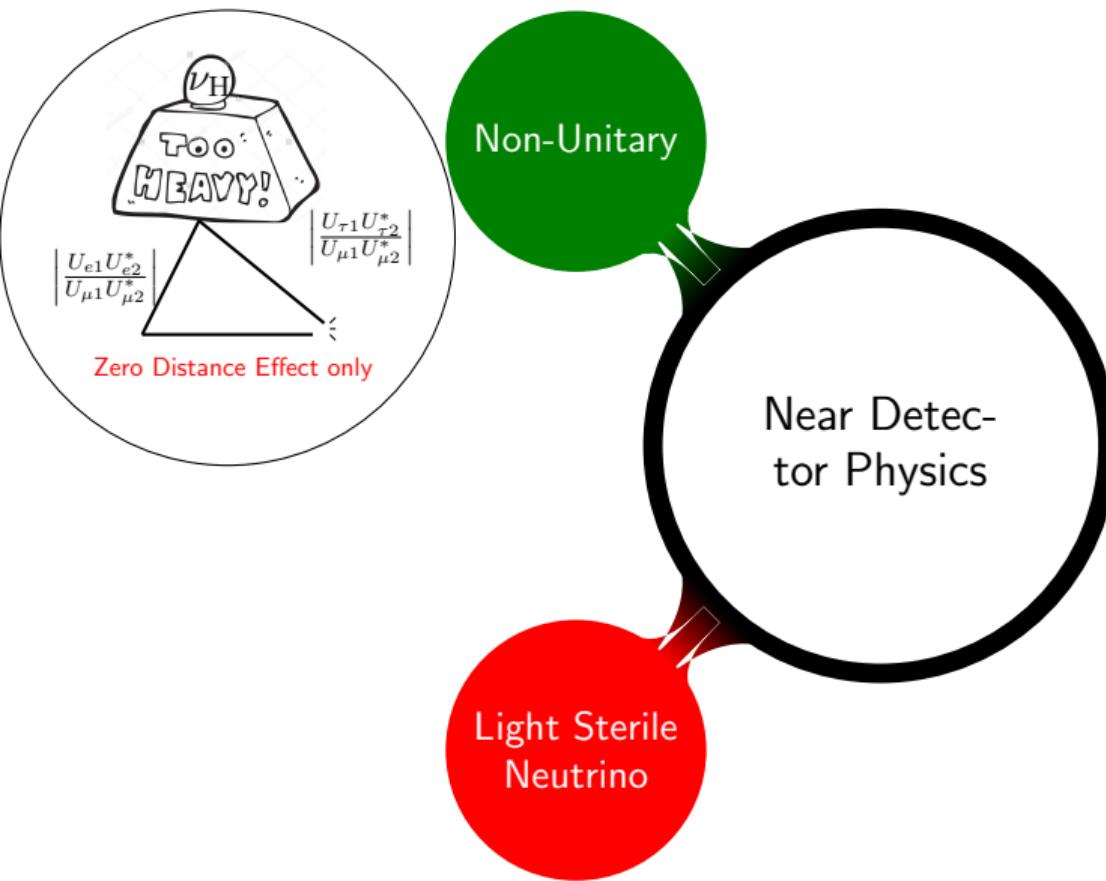
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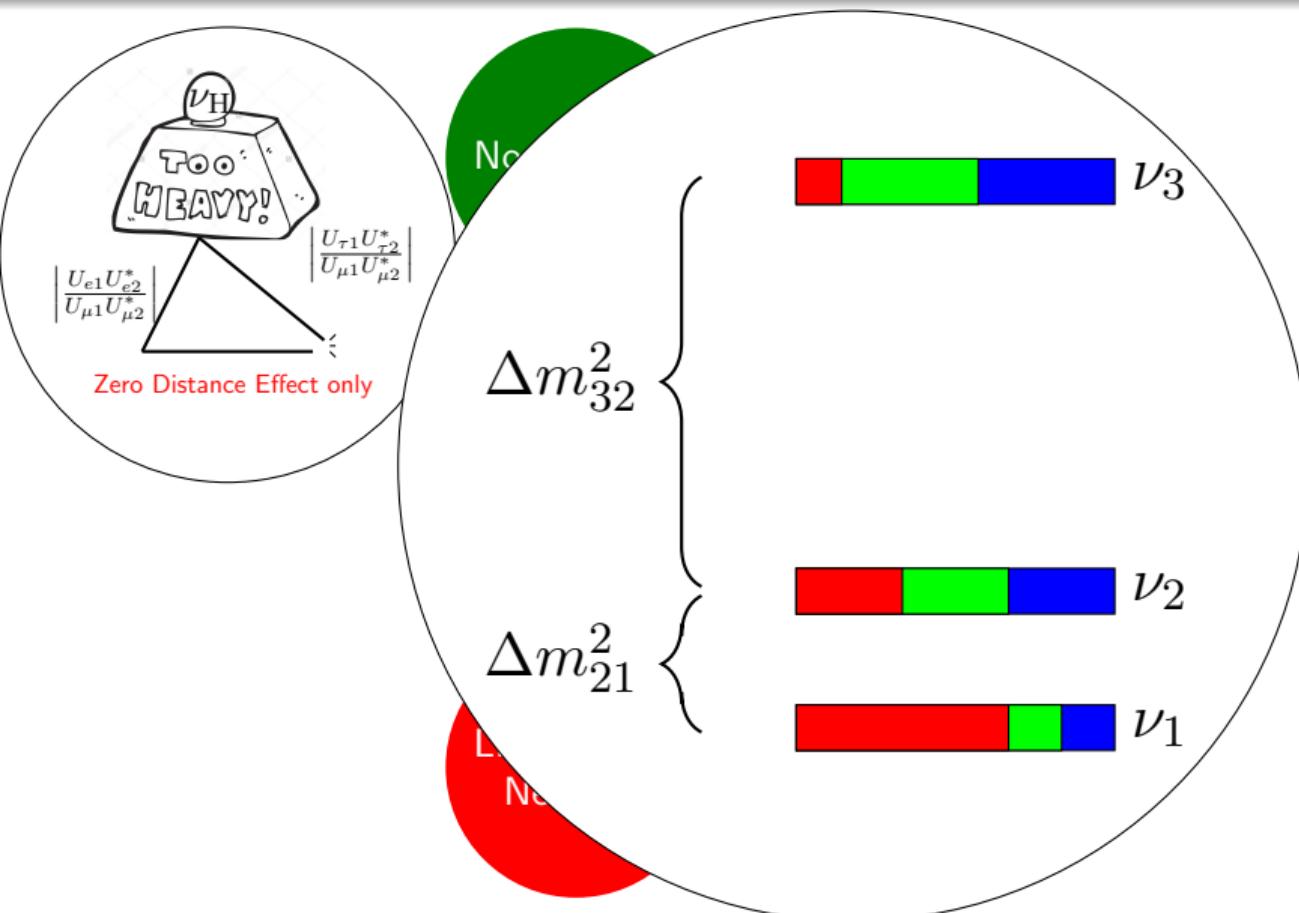
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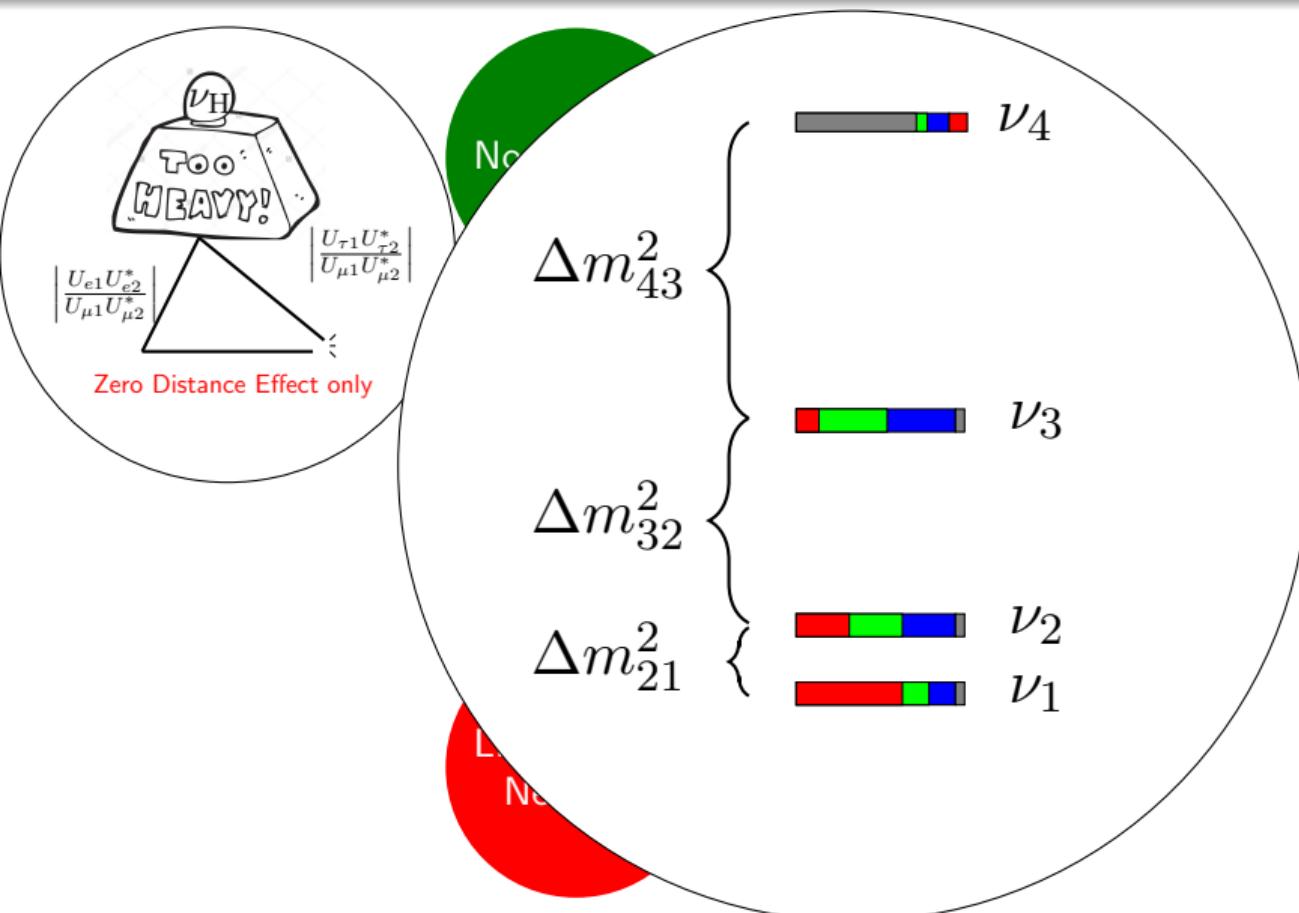
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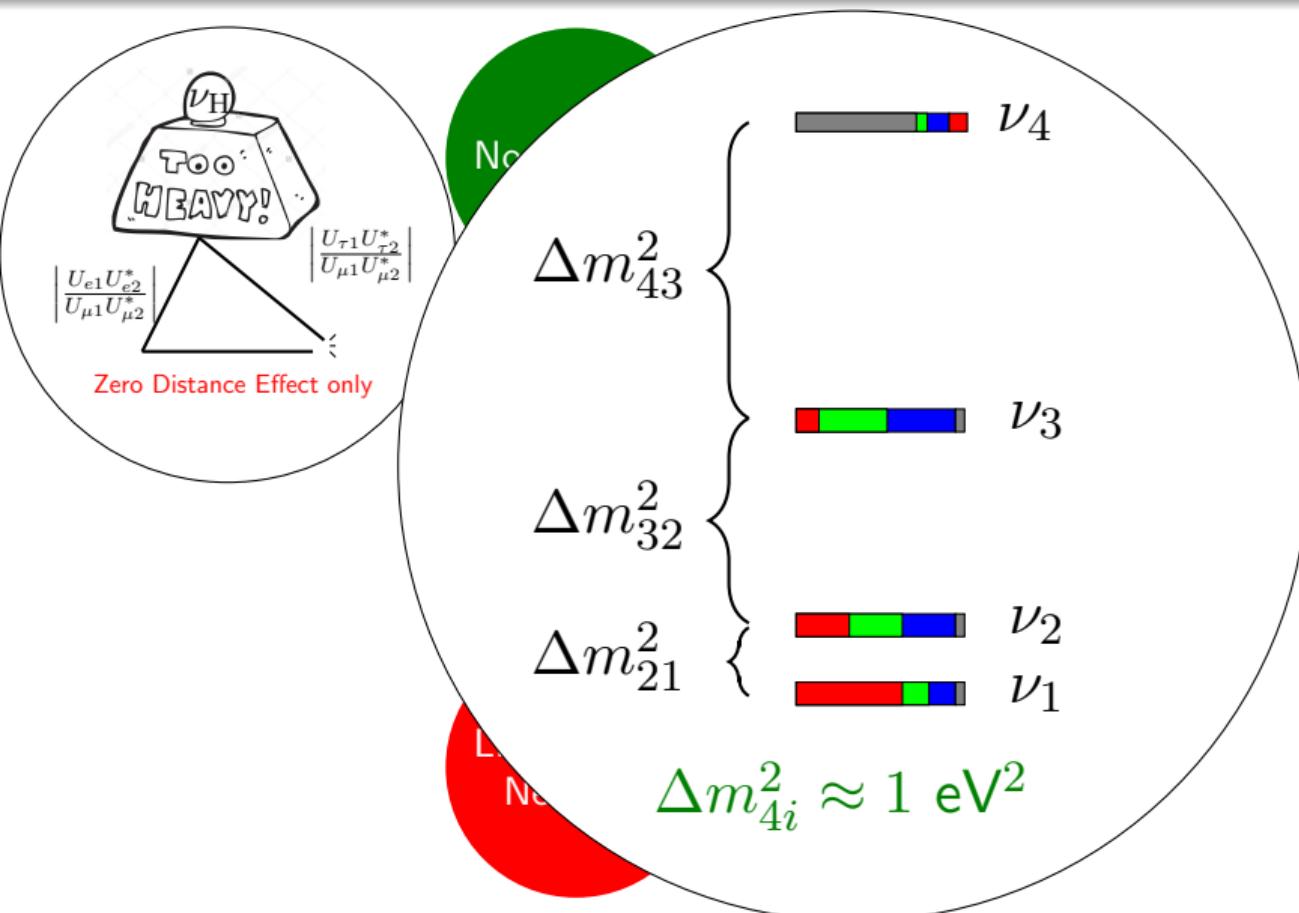
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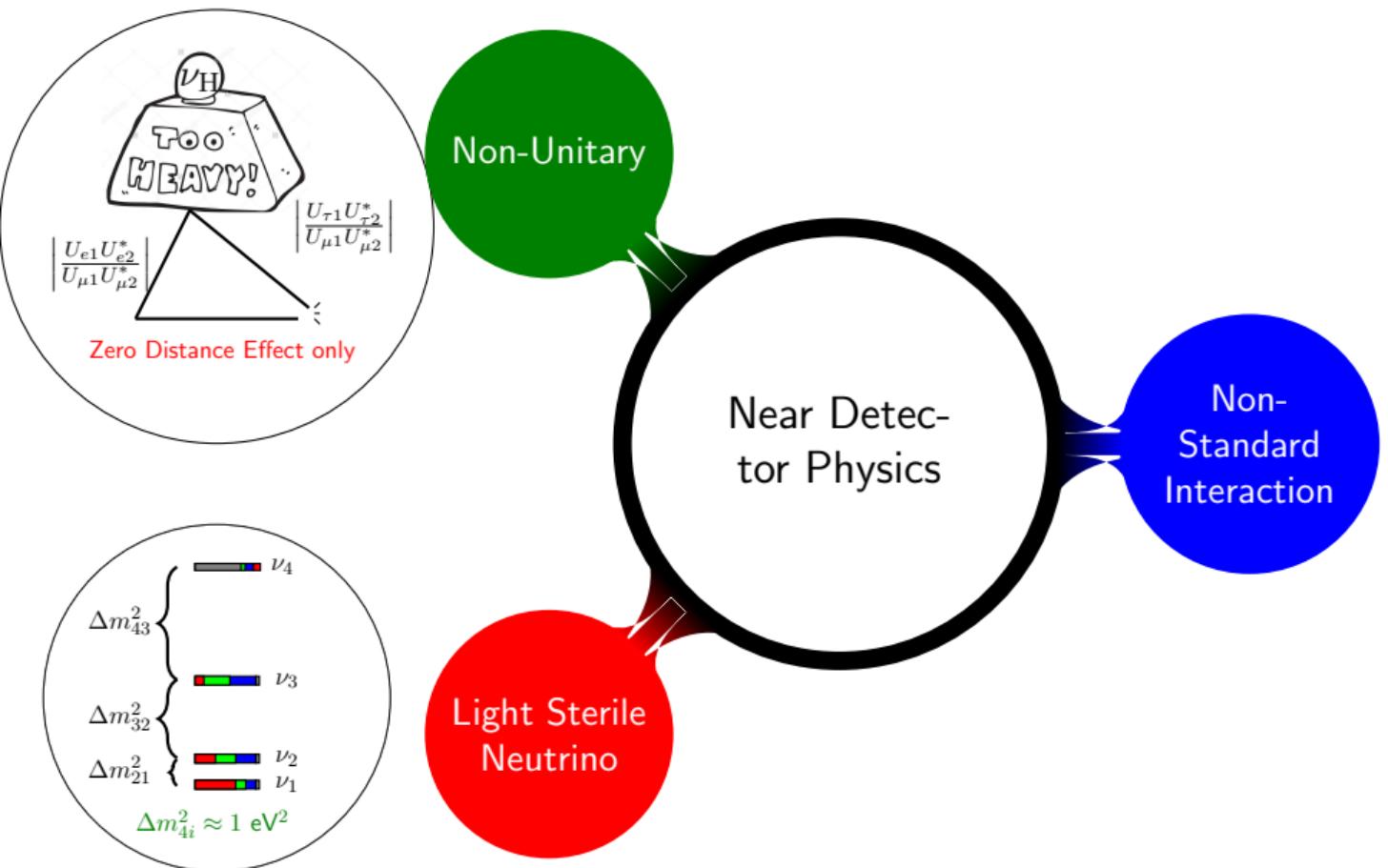
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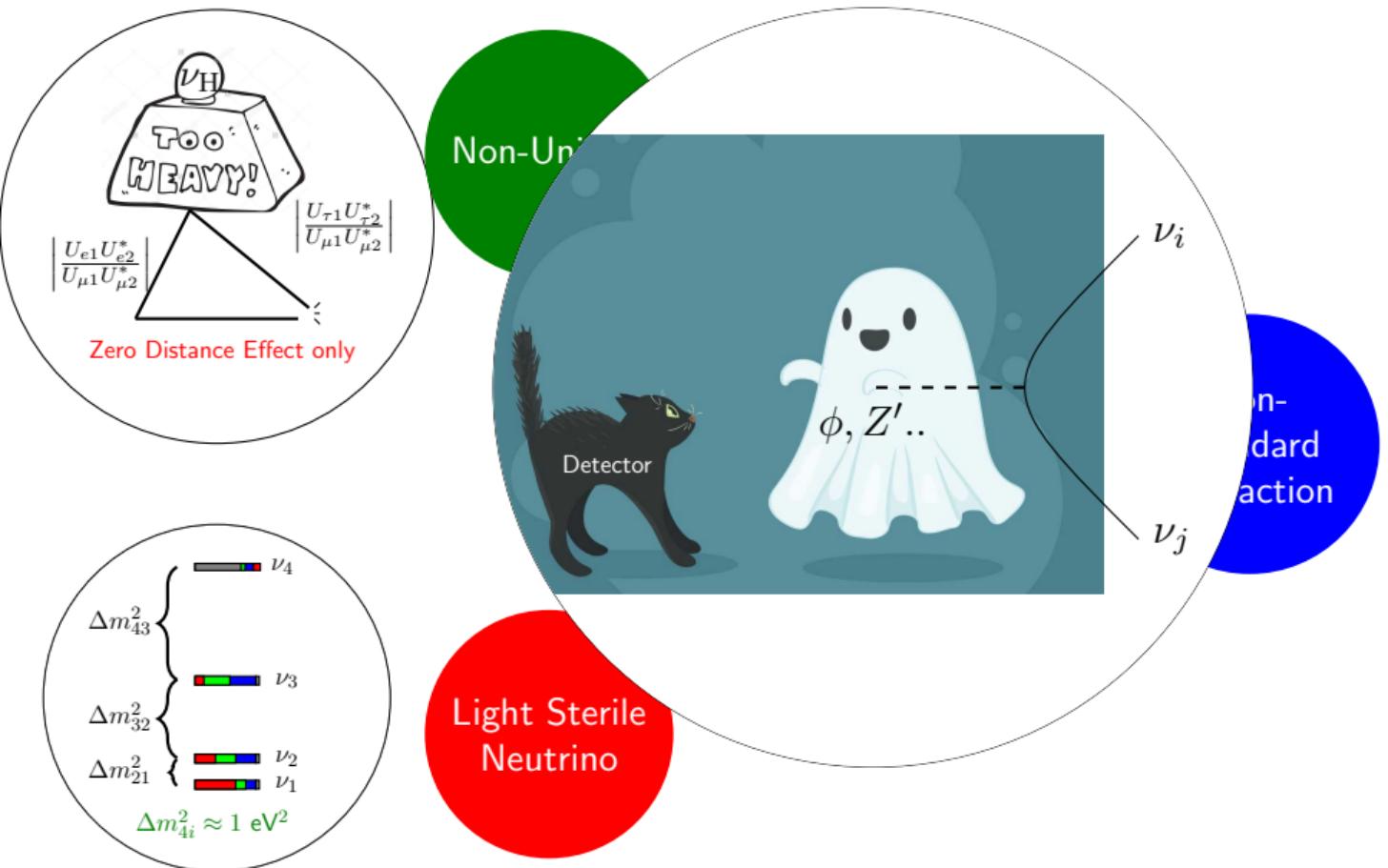
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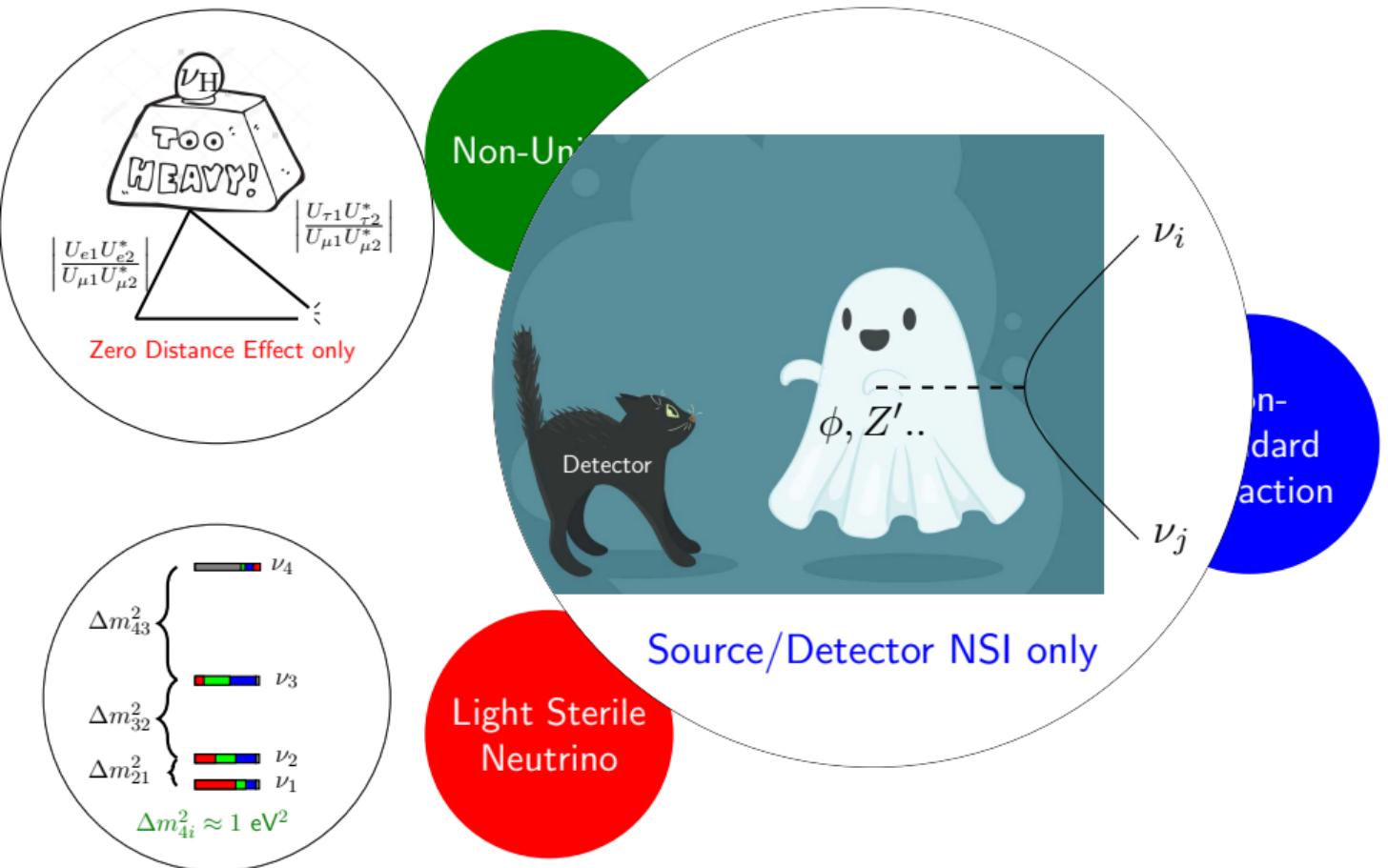
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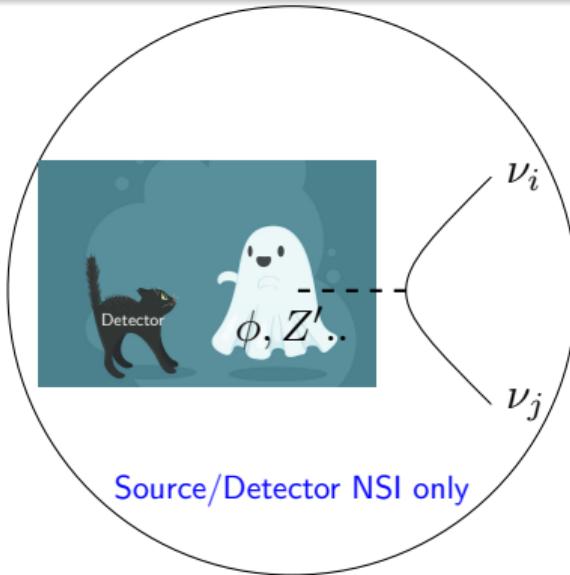
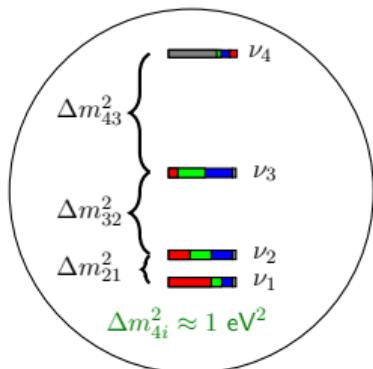
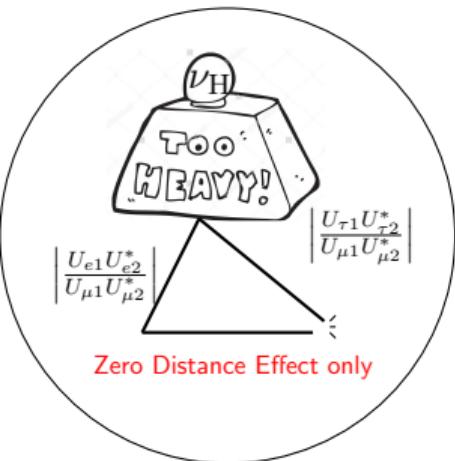
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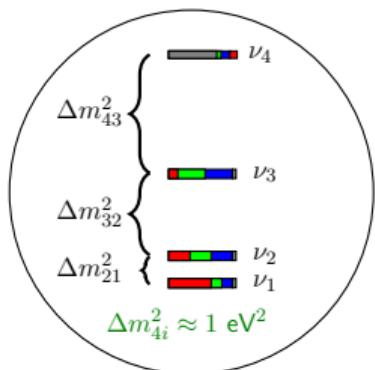
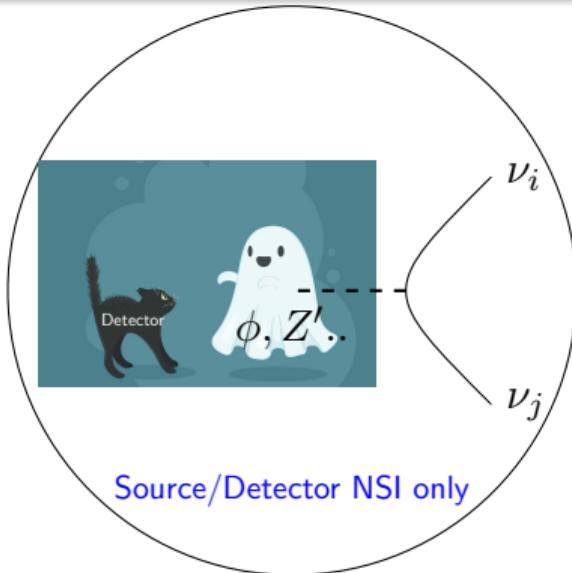
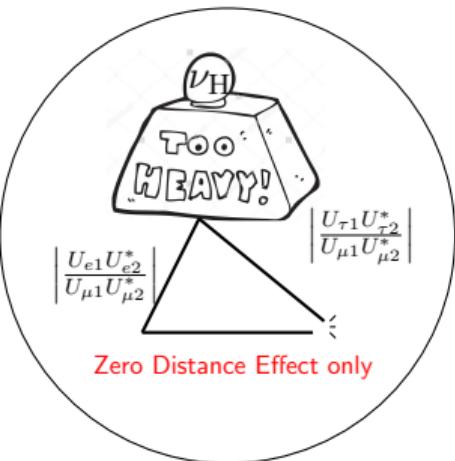
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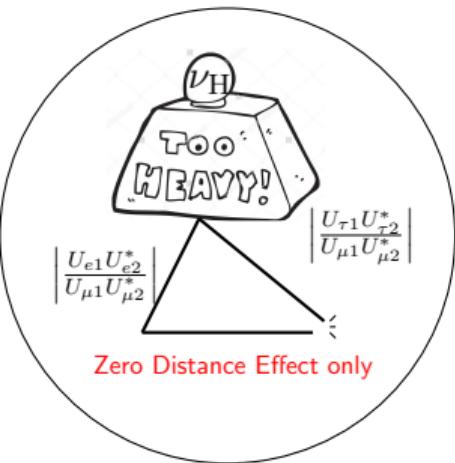
New physics change ν_e spectrum



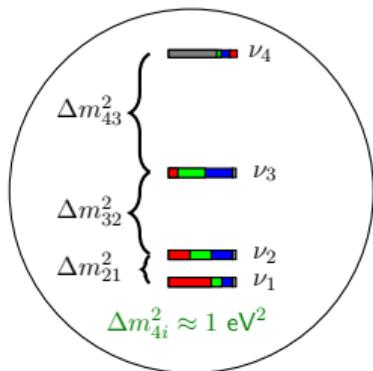
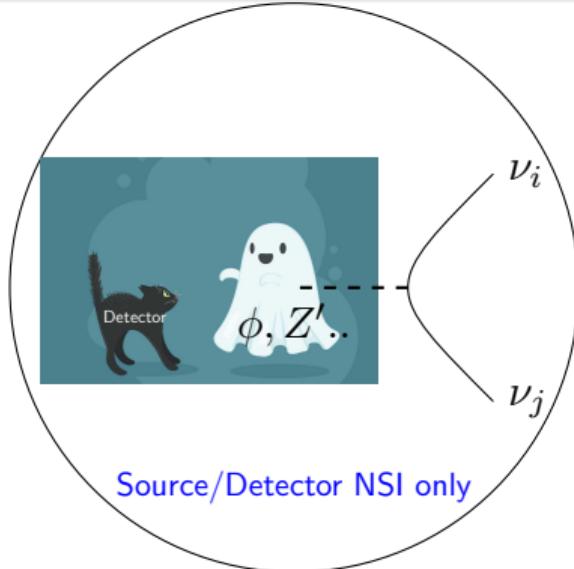
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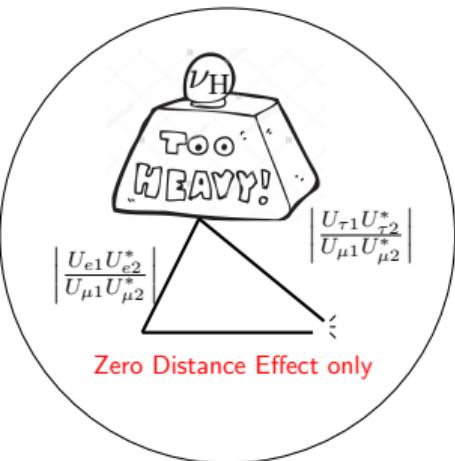
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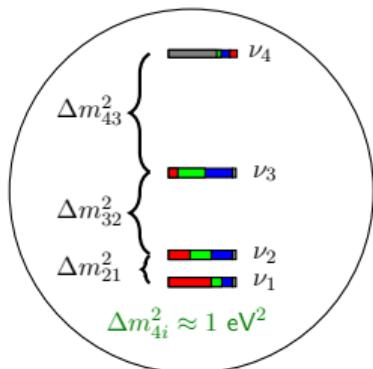
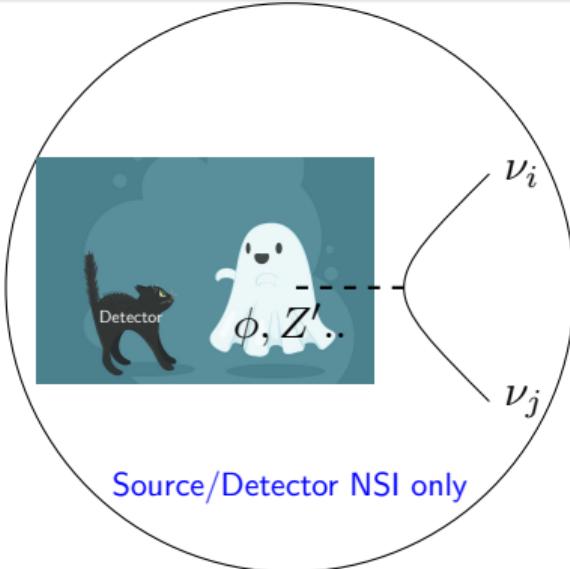
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$$P(\nu_\mu \rightarrow \nu_e) = 1 - \sin^2 2\theta_{\mu e} \sin \frac{\Delta m_{41}^2 L}{4E}$$

We can put better constrain to new physics!

We simulated:

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SBNE = SBND + μ BooNE + ICARUS

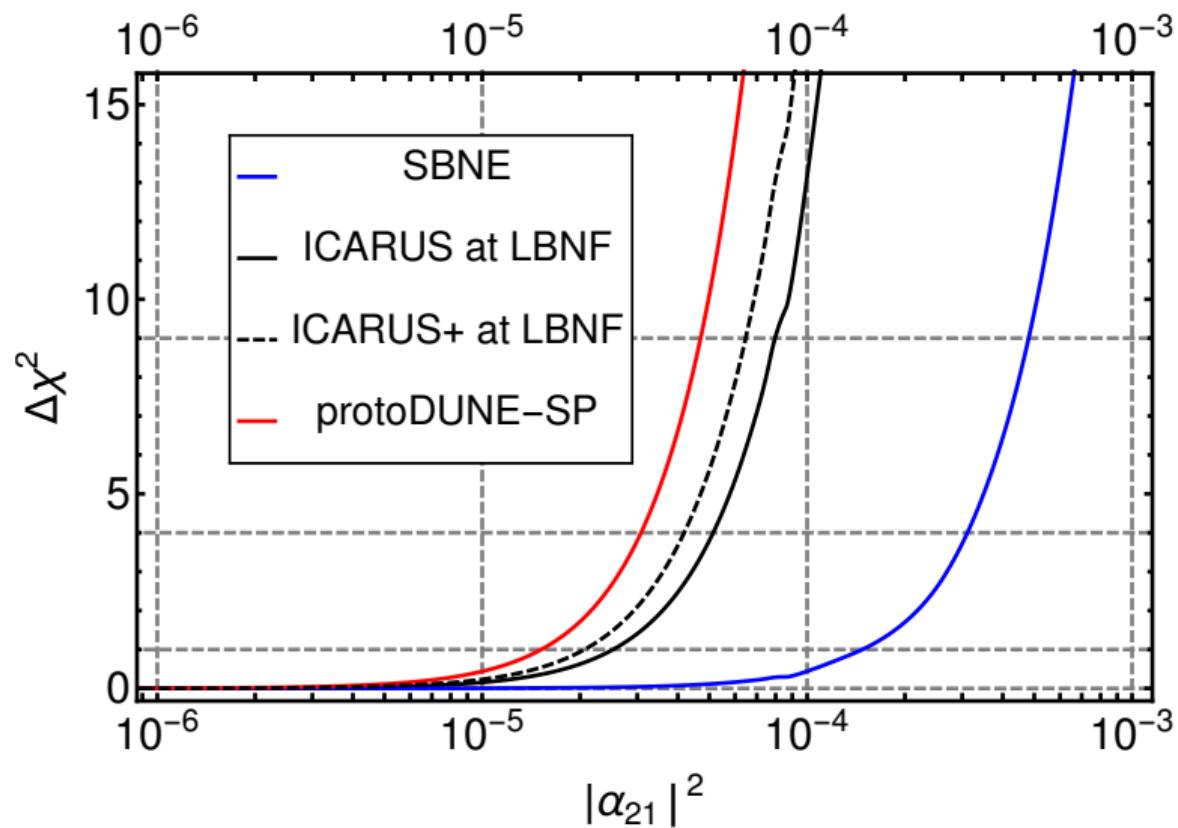
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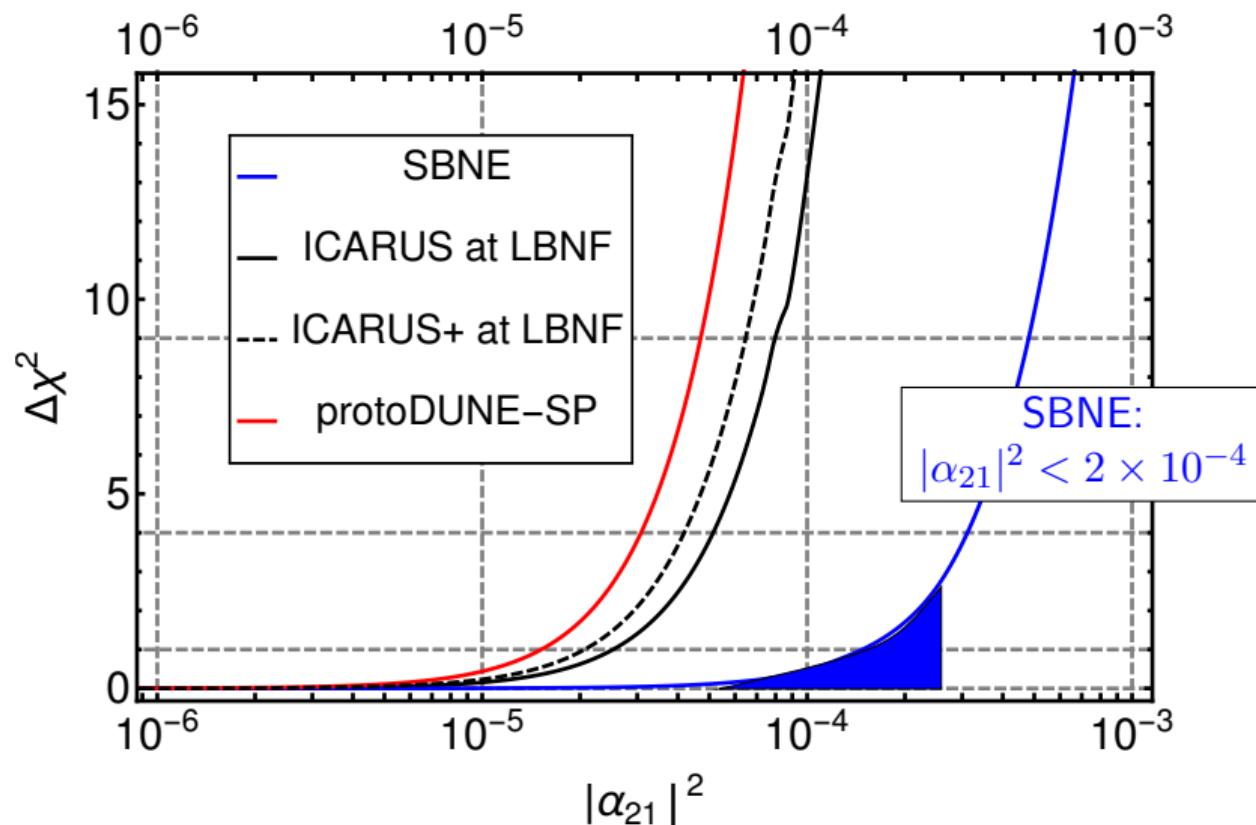
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LBNF beam with: protoDUNE and ICARUS as ND

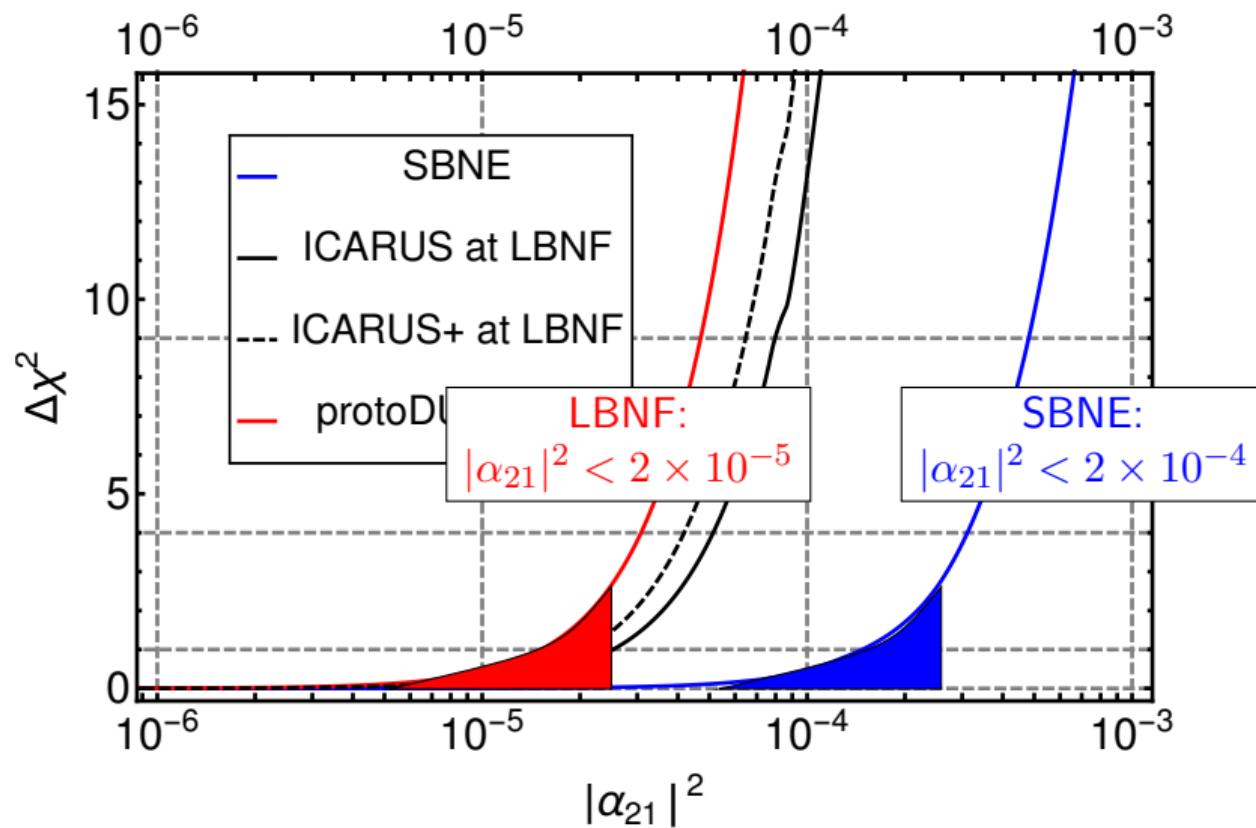
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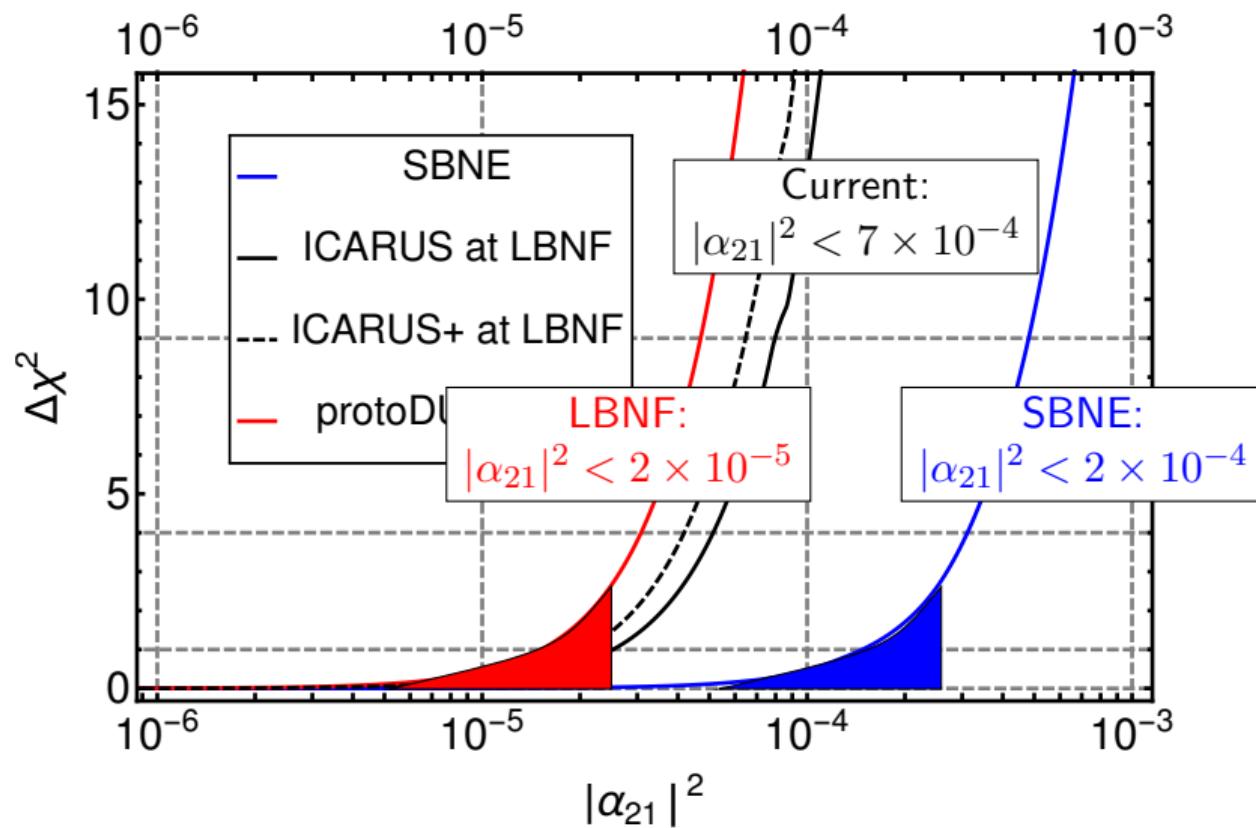
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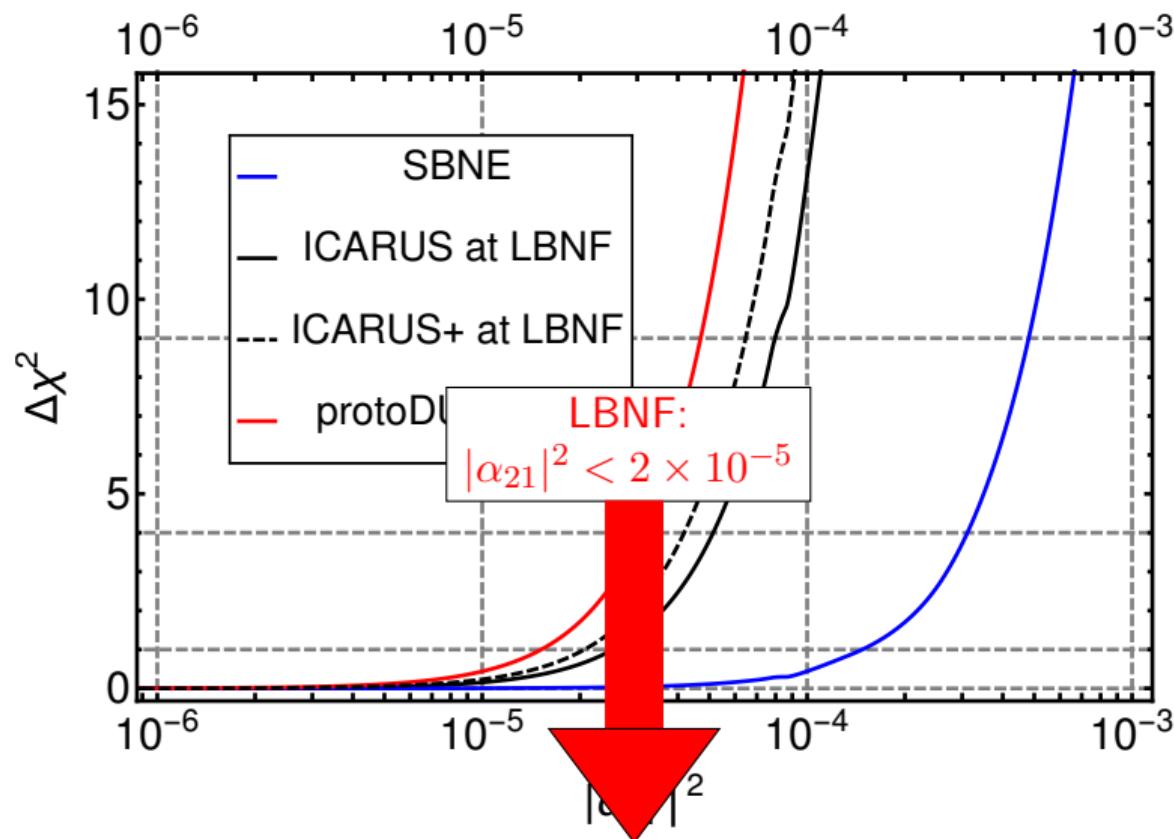
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Can we really reach this level?

We need to know the expected flux precisely!

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Thus, we need to know the expected flux:

We need good production simulation and measurement (μ , hadron)

We parametrized our lack of knowledge

Let's parametrize our lack of knowledge to see its impact:

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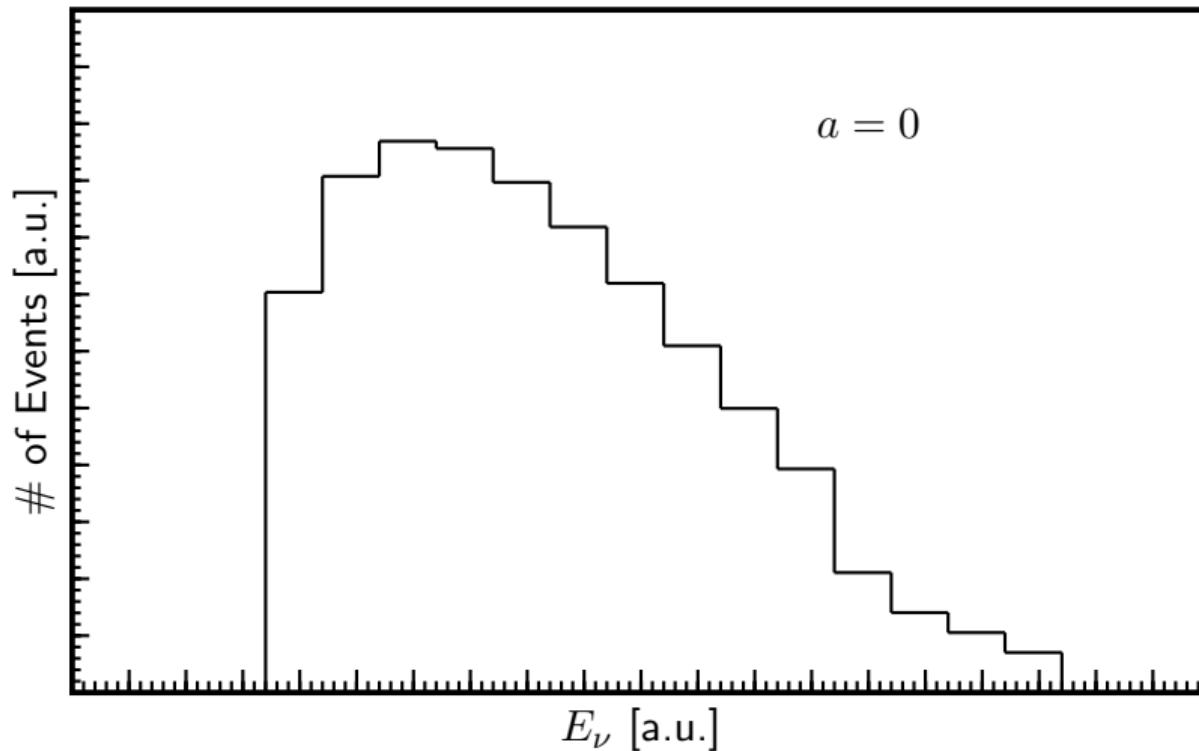
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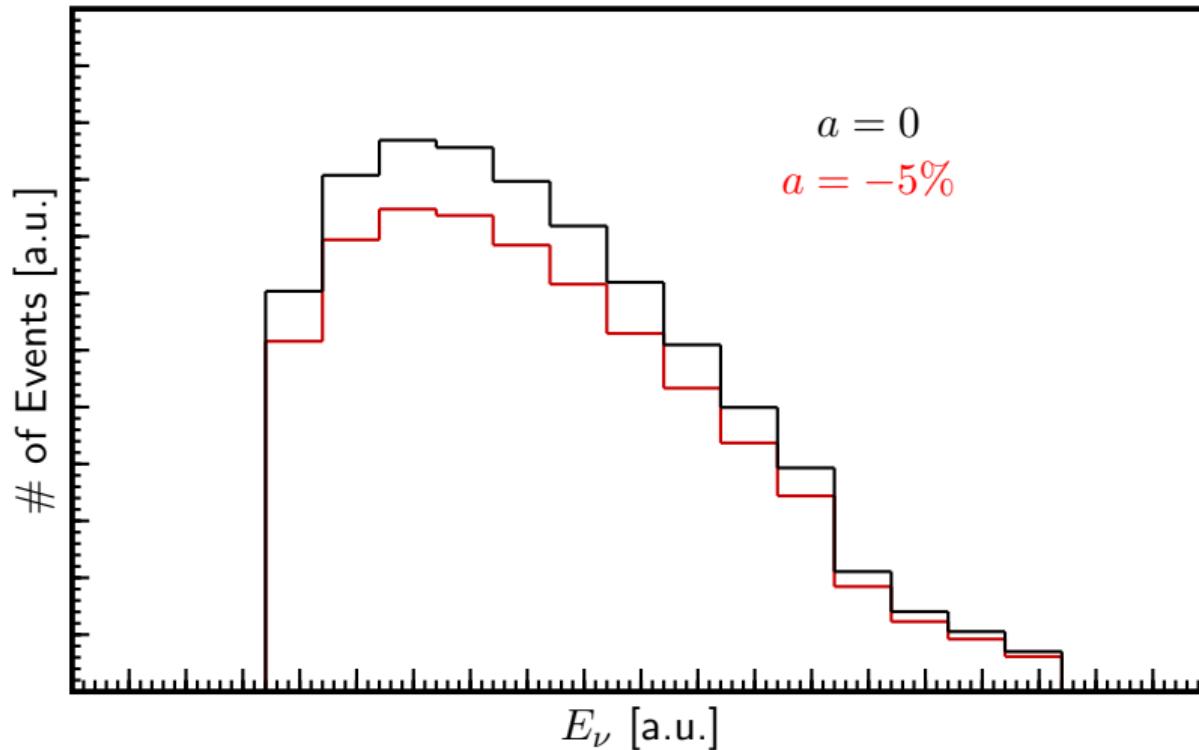
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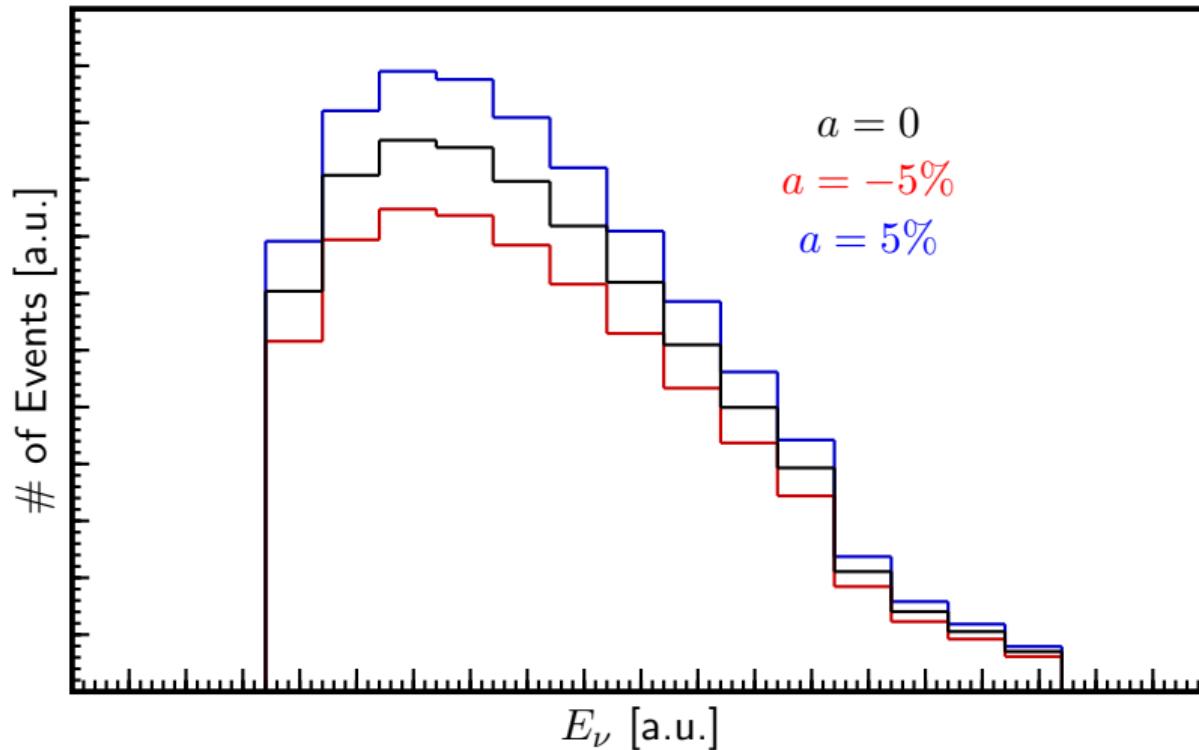
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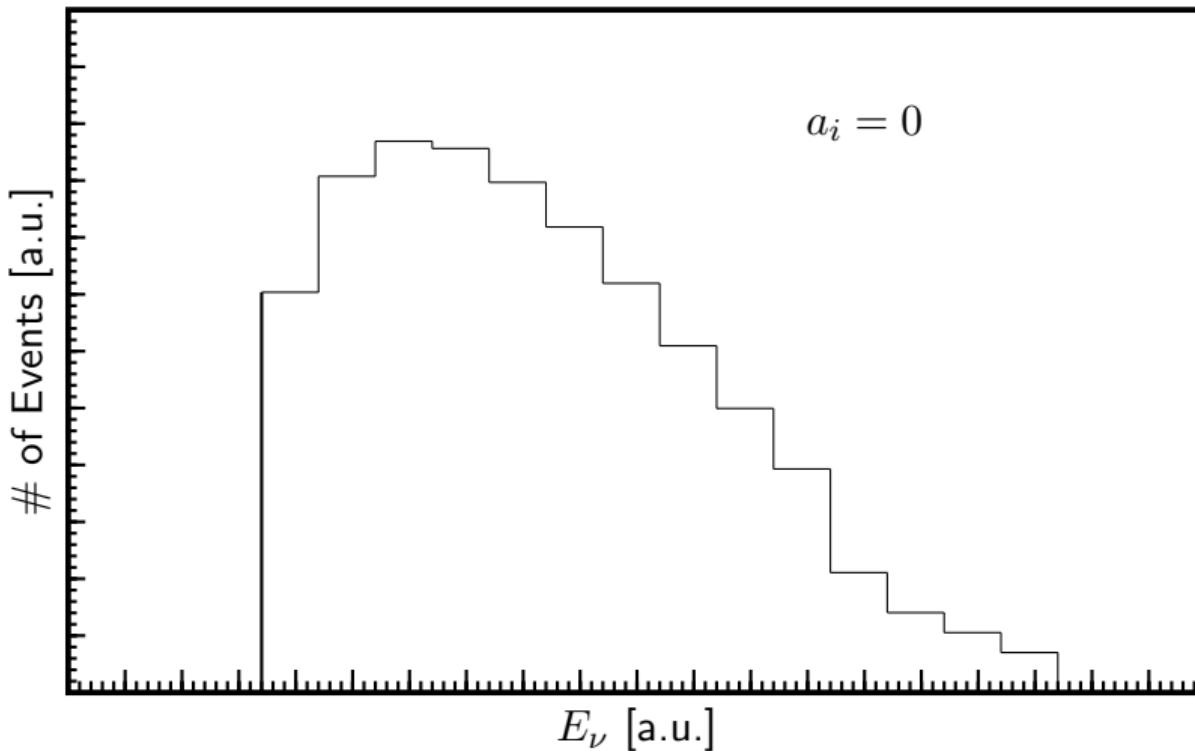
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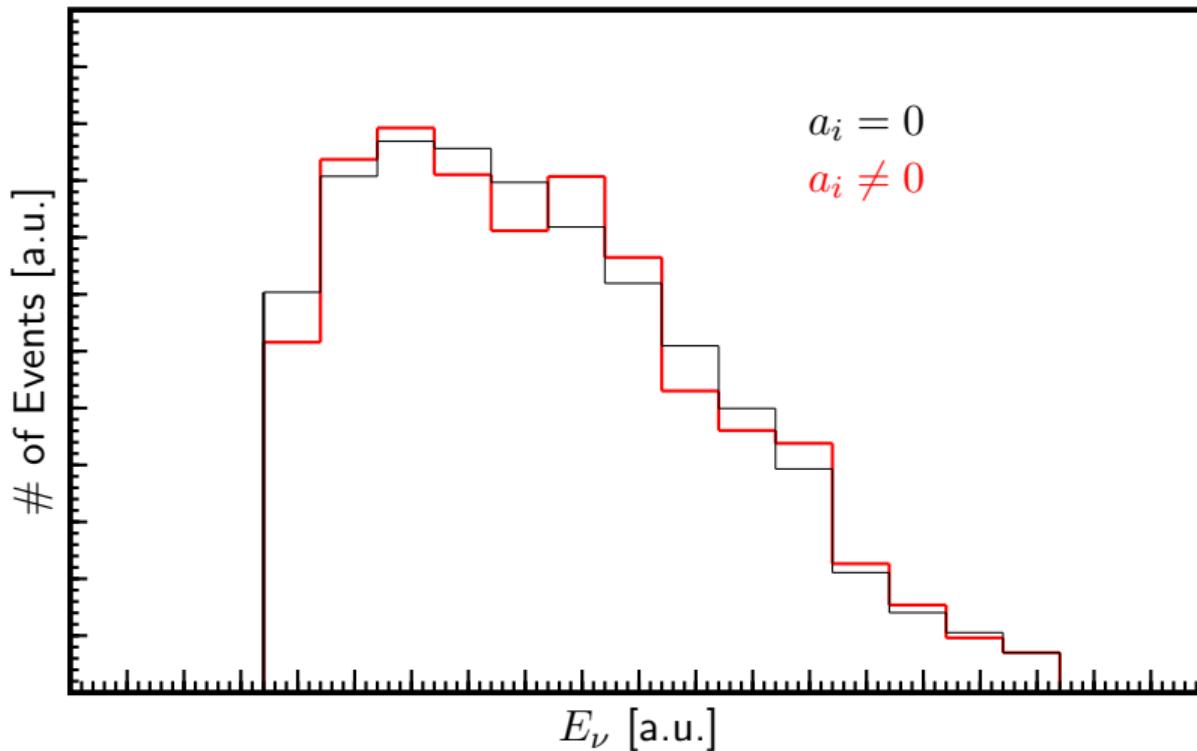
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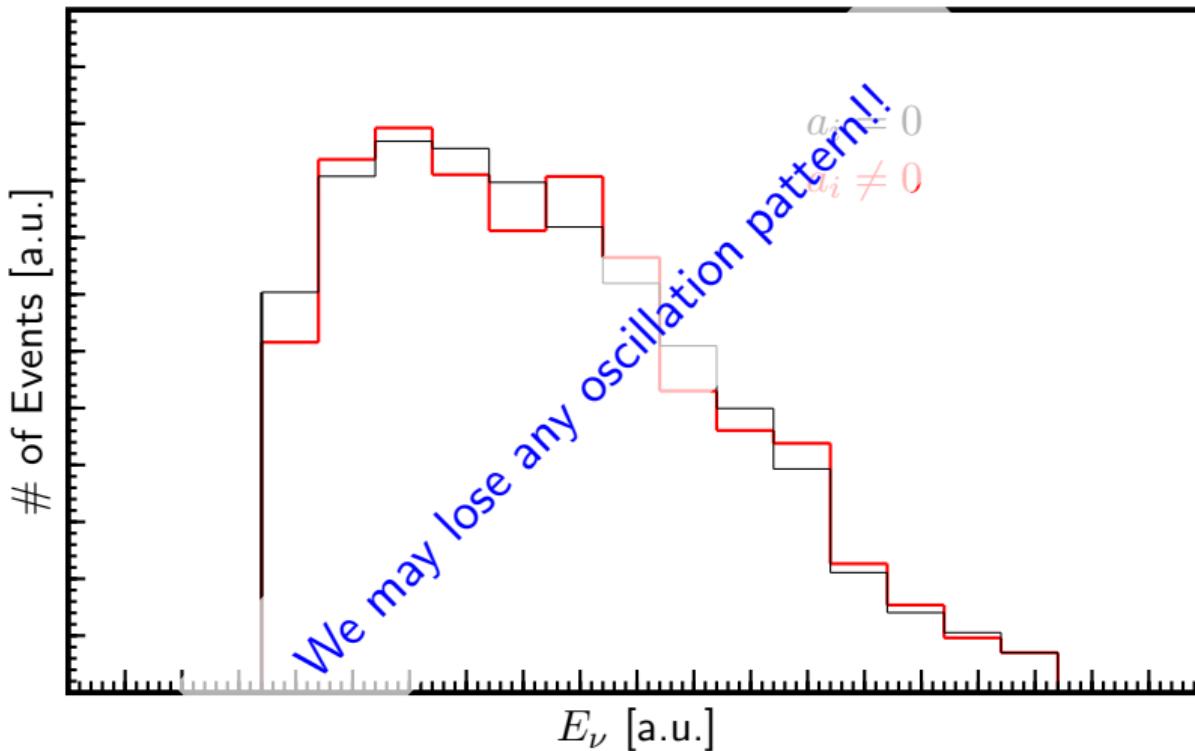
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σ_s the real parameter here

A bit of math....

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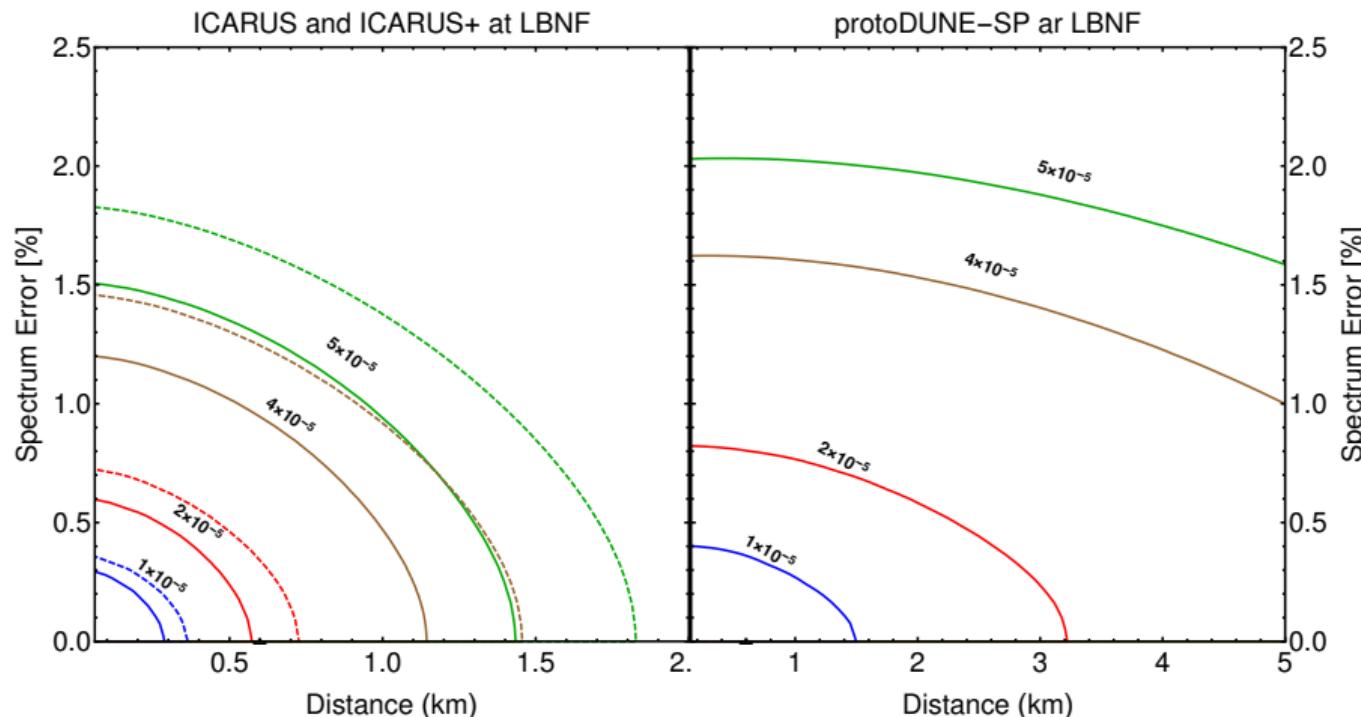
$\sigma_{sa} = \sigma_{sb} = \sigma_s$ Spectrum error

We need $\sigma_s \sim O(1)\%$

What we got (for $|\alpha_{21}^2|$):

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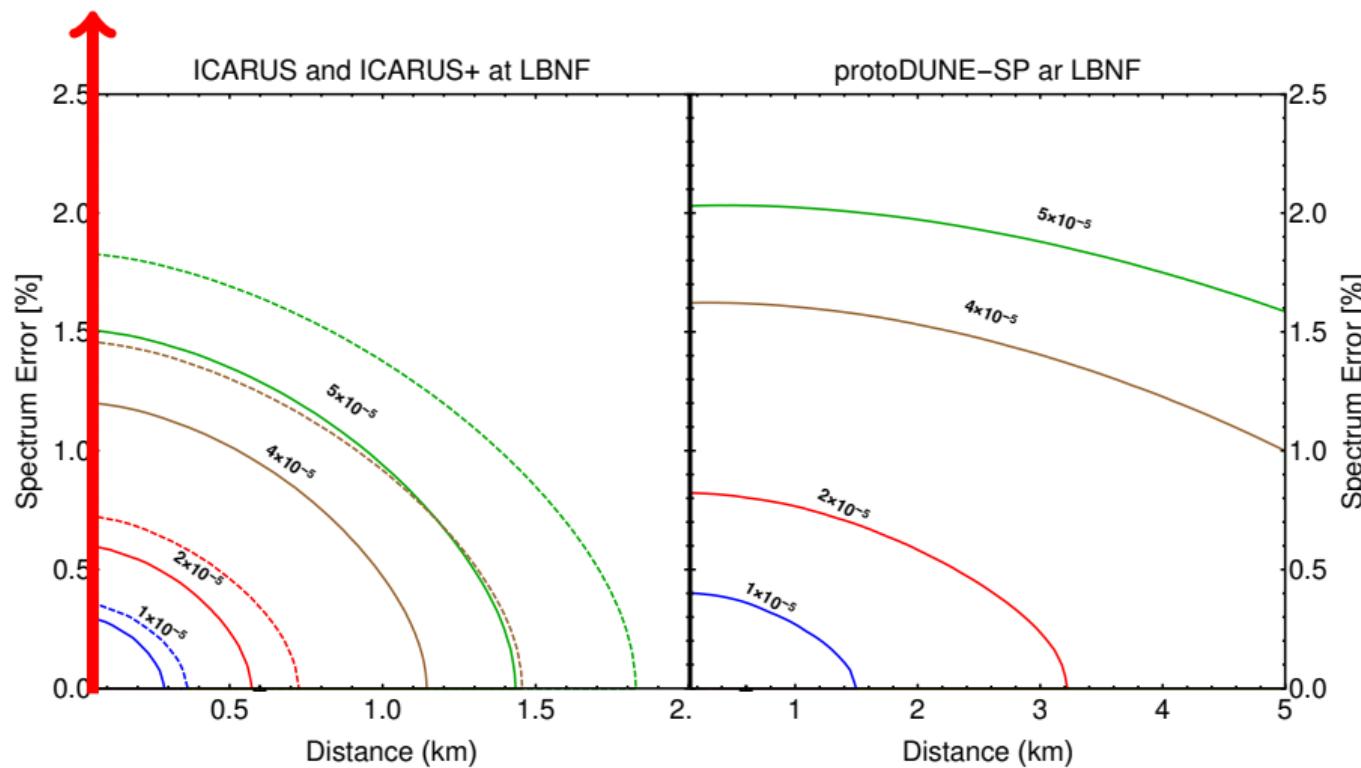
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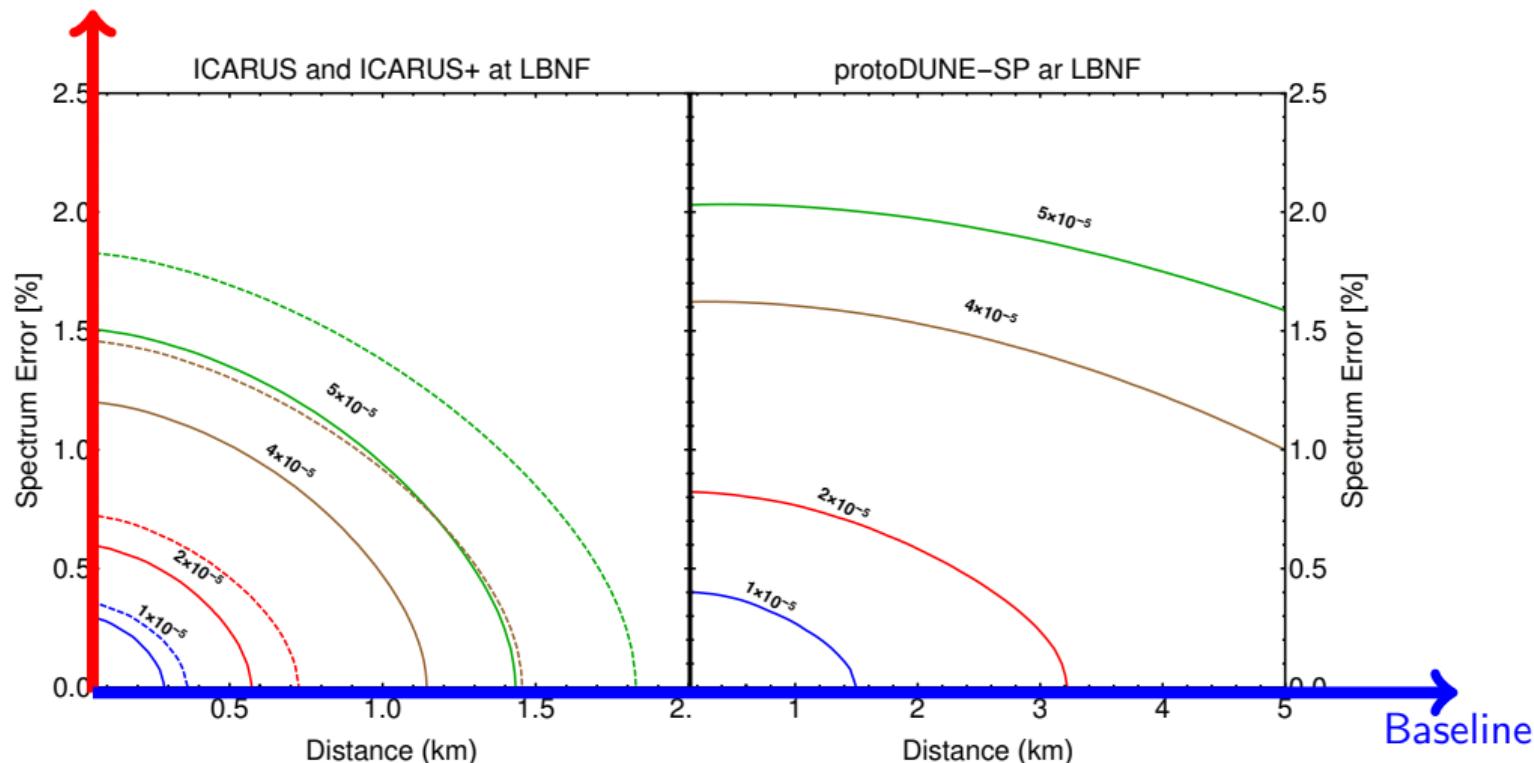
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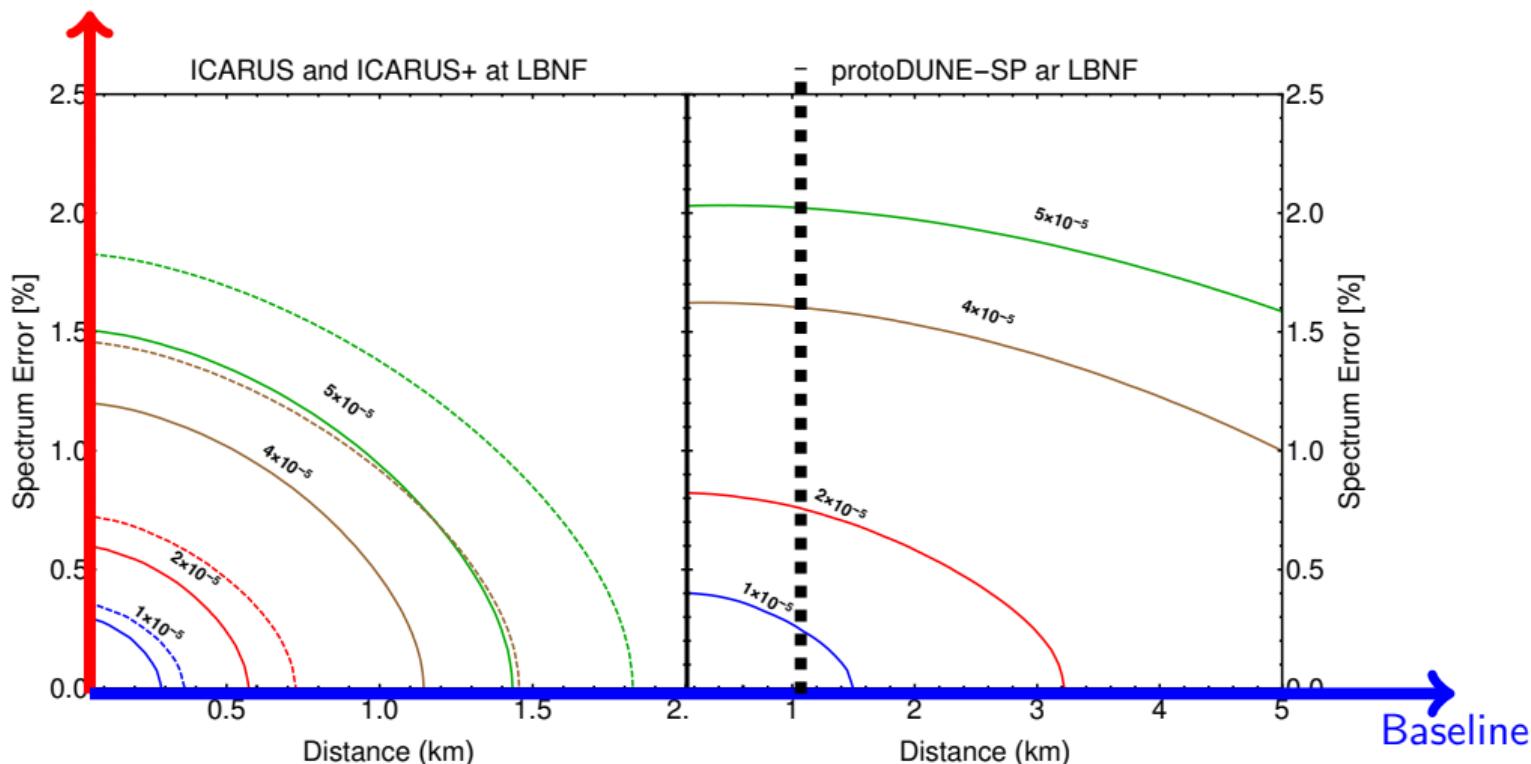
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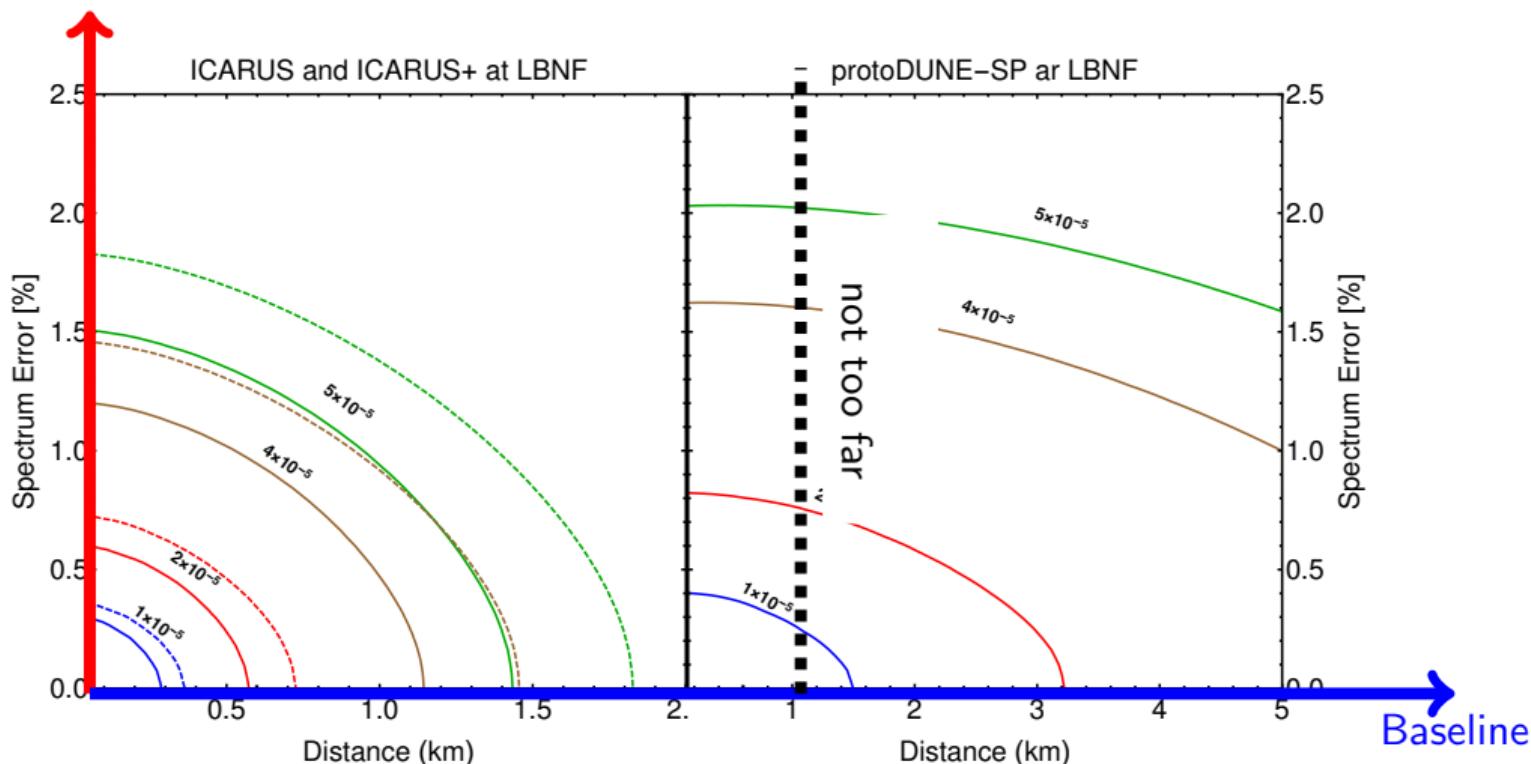
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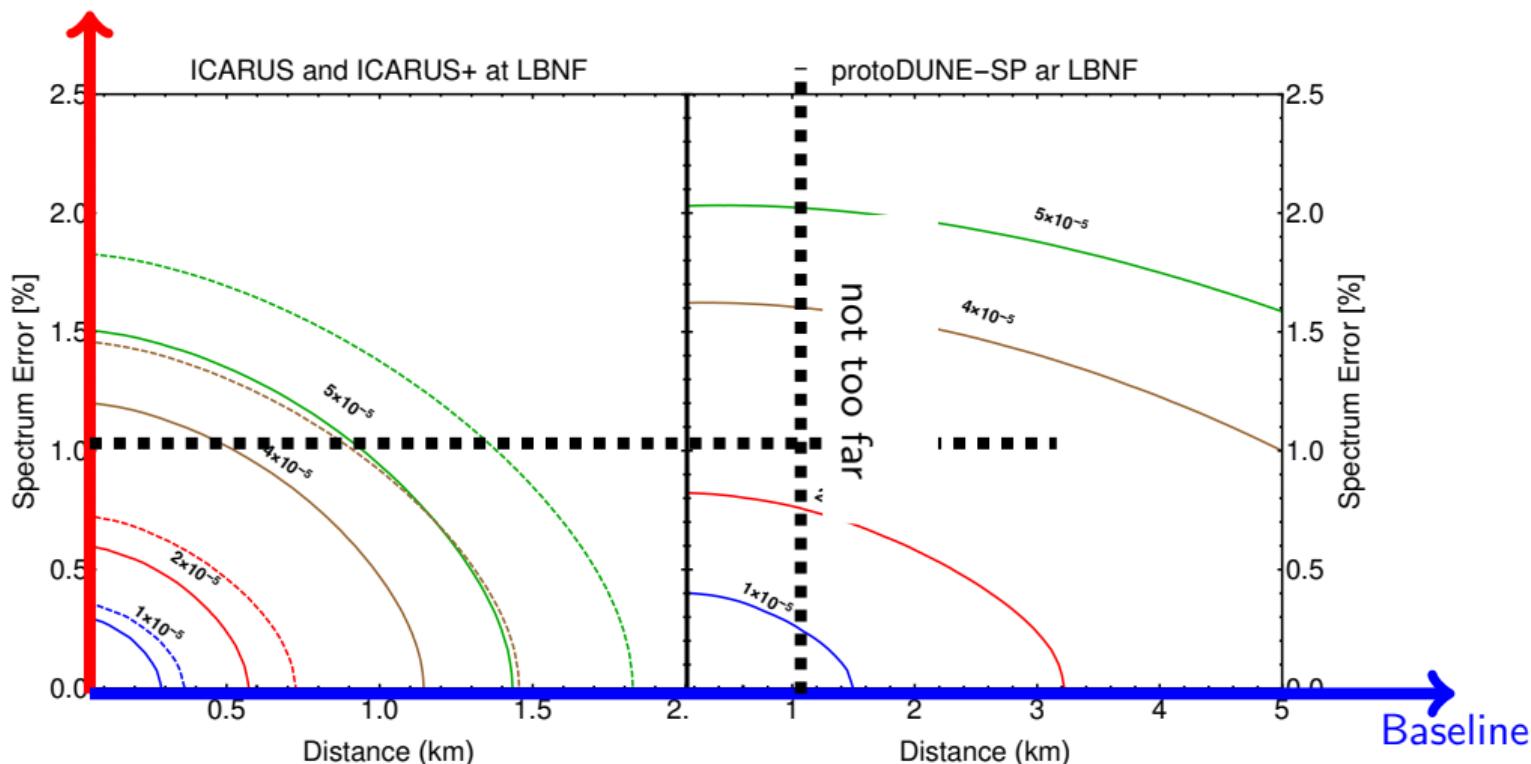
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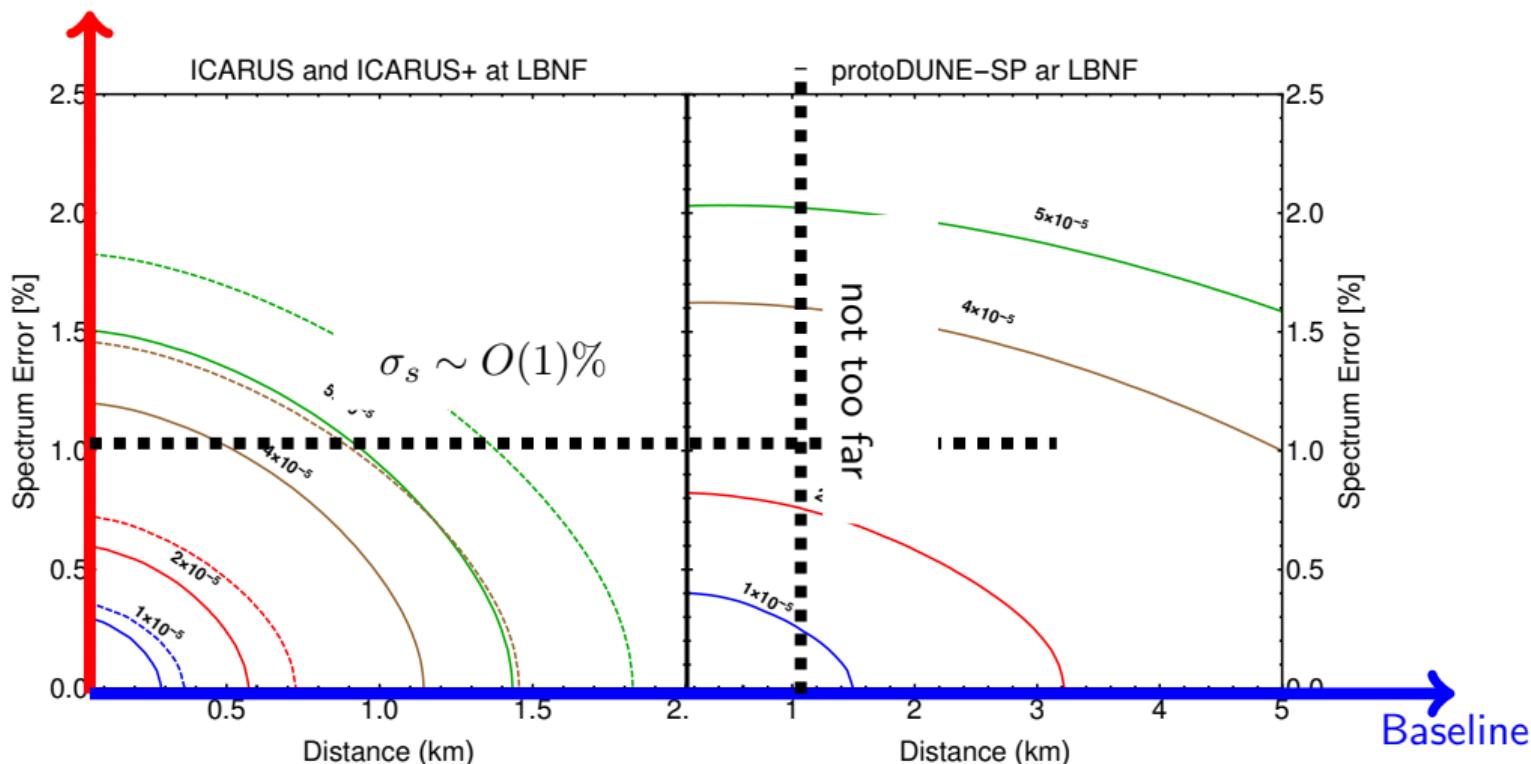
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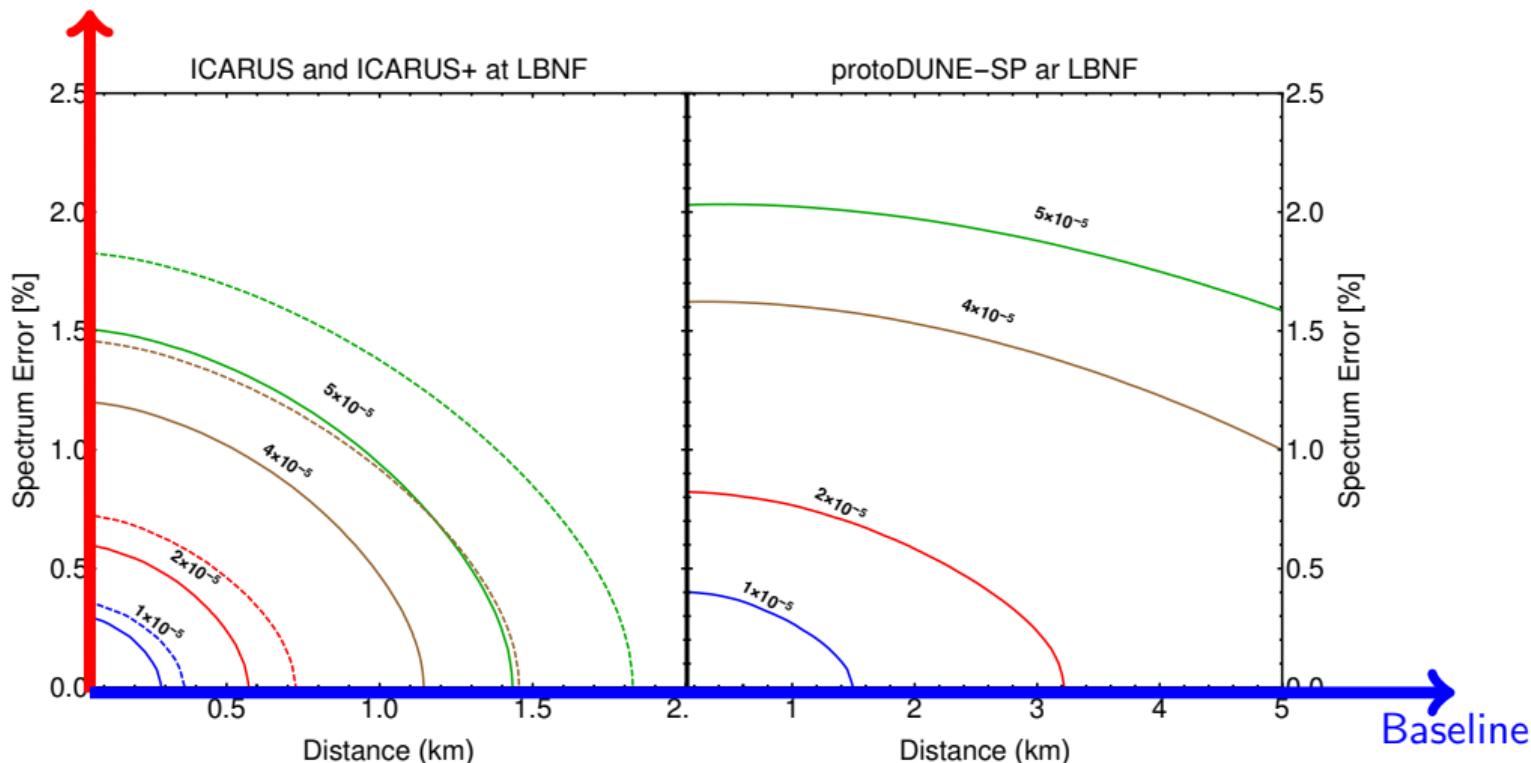
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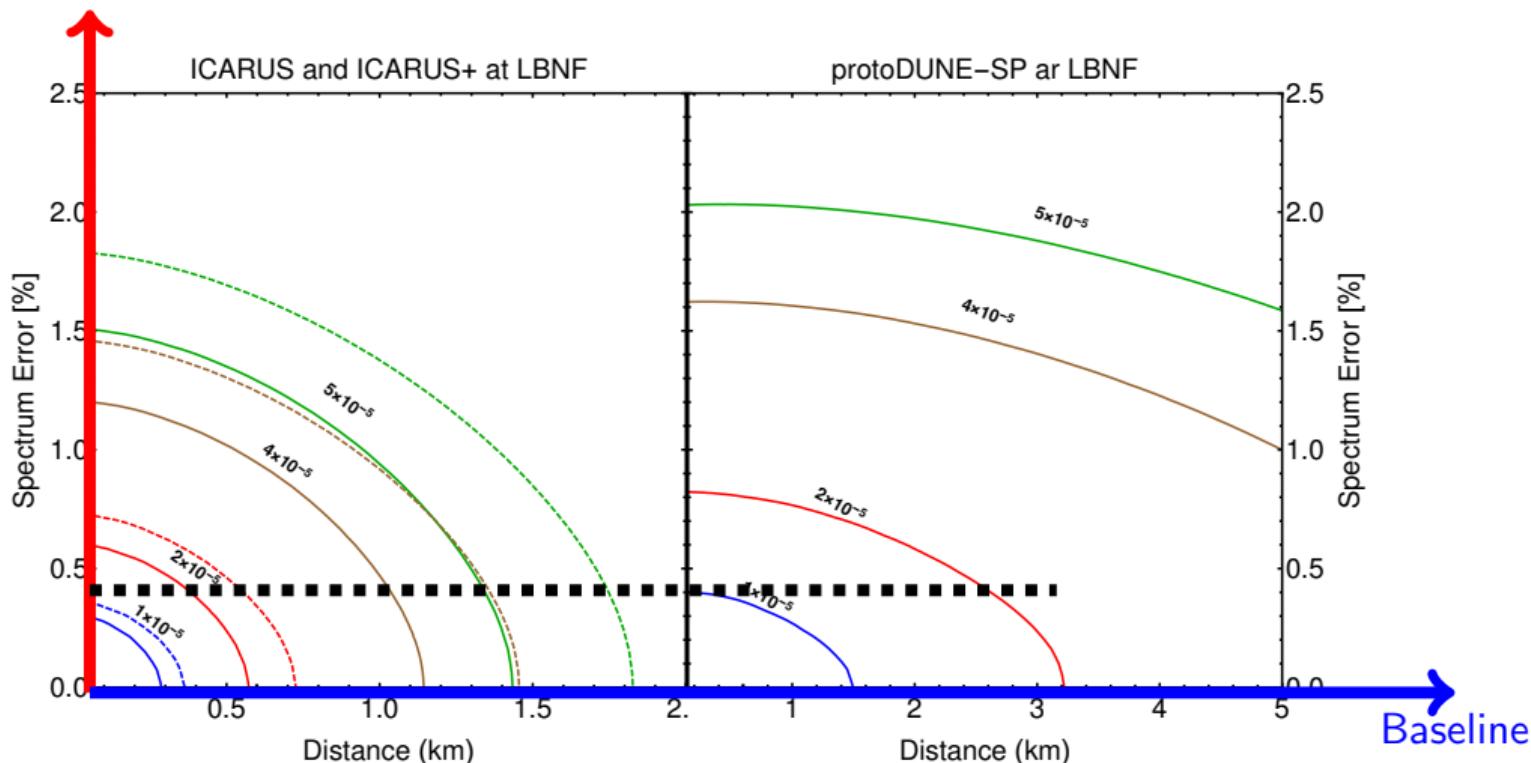
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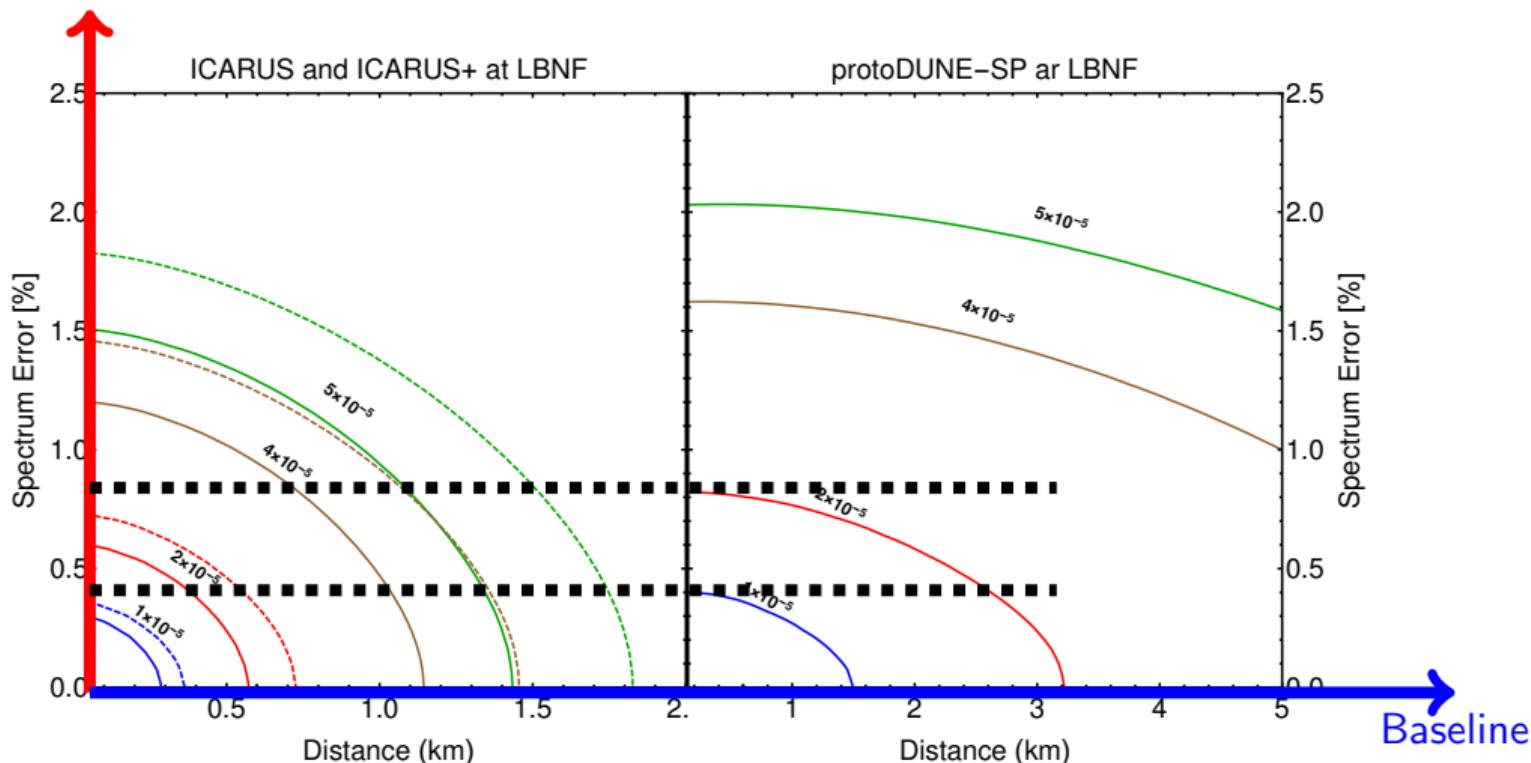
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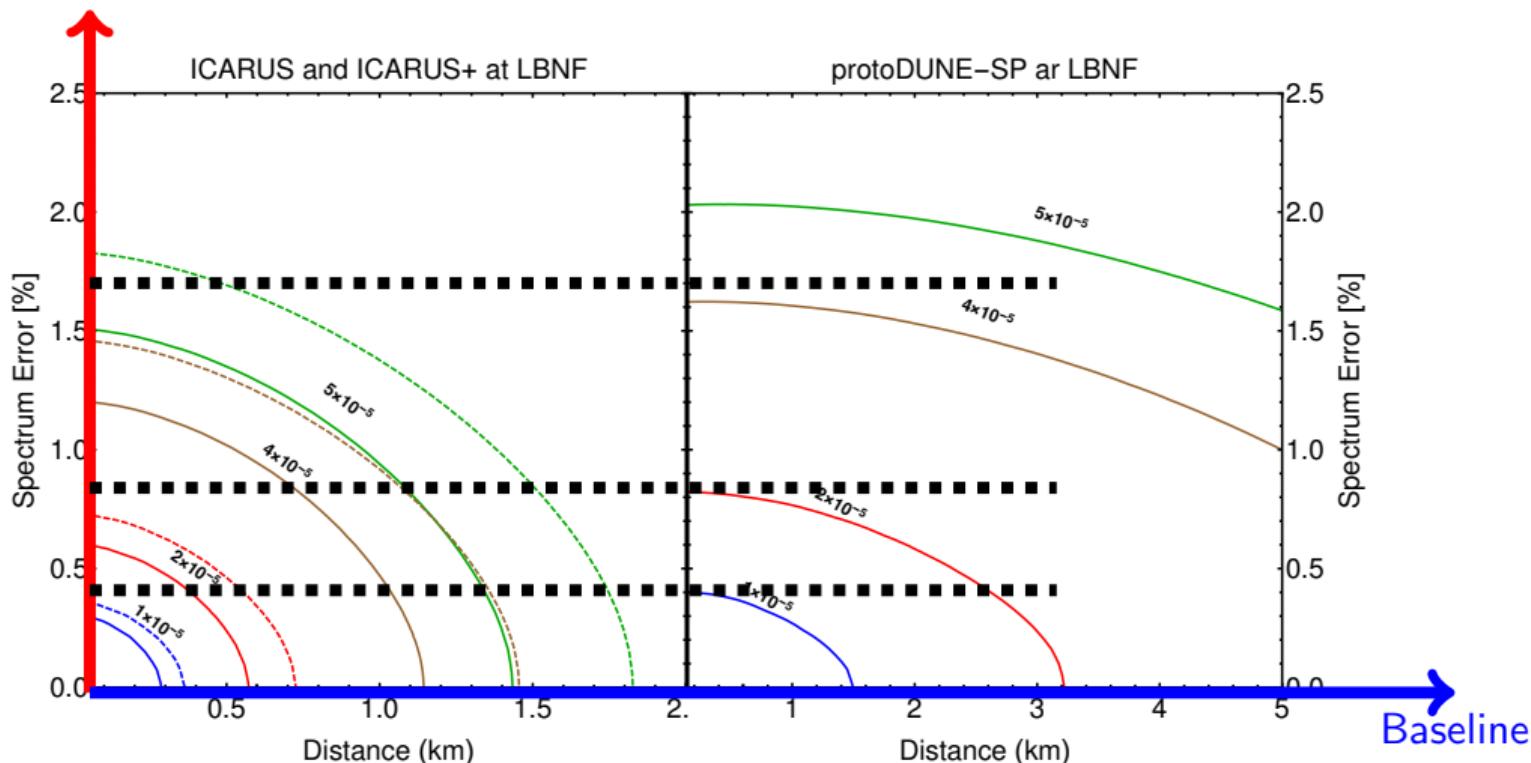
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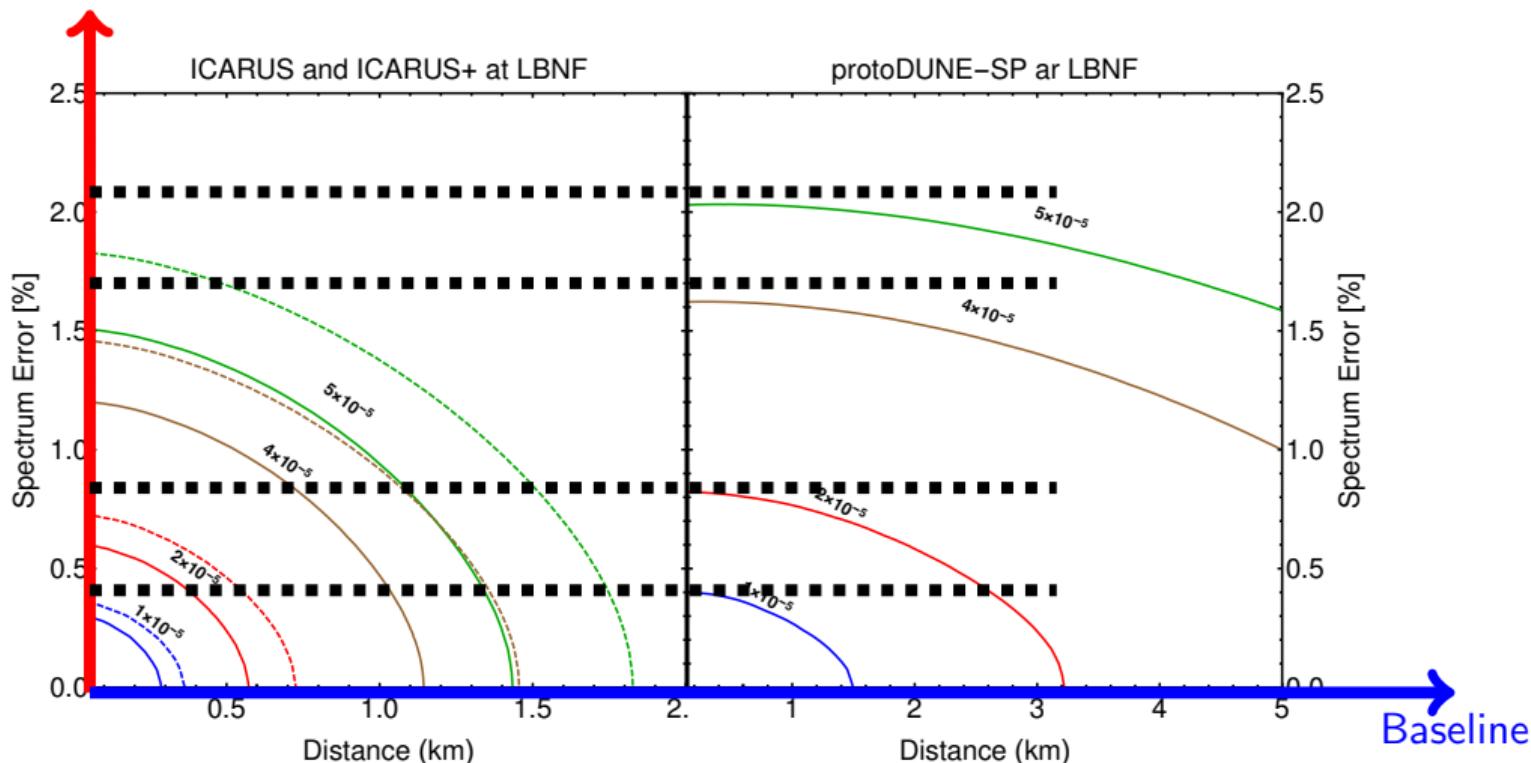
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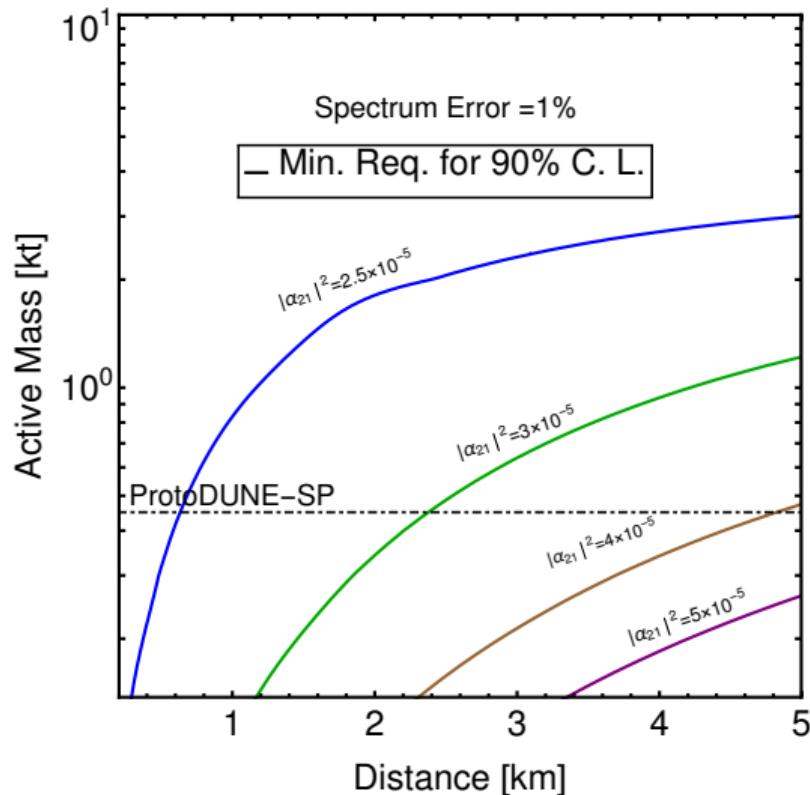
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Setting a σ_s goal, we can get minimum requirements

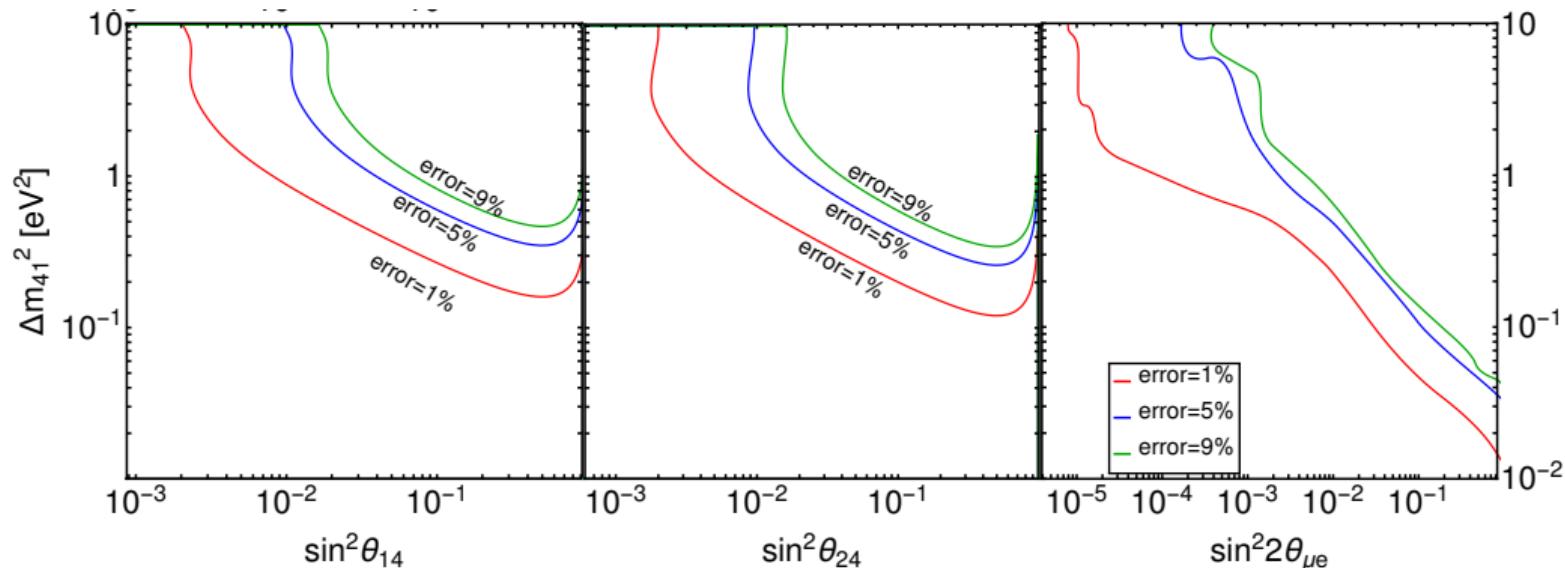


We can probe sterile neutrinos too!

similar for sterile neutrino!

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Thanks for the supporters

Thanks

G. V. Stenico for SBN codes



Generalitat Valenciana

Ramón y Cajal

CONACyT and SNI (Mexico).